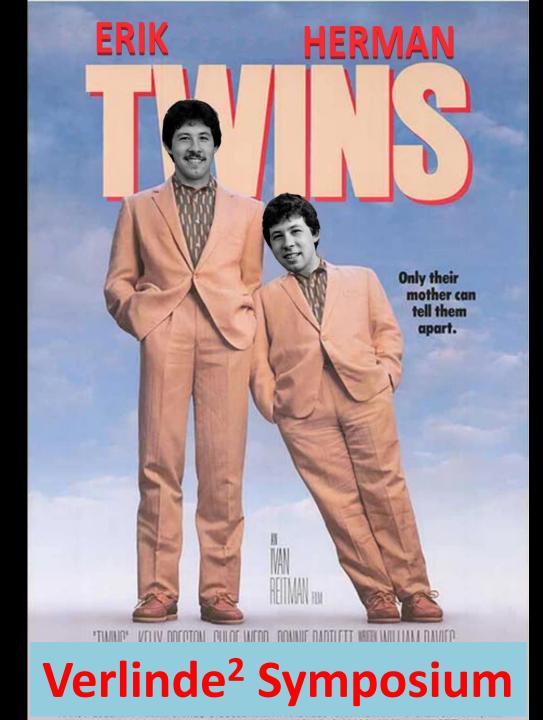
Verlinde² Symposium





THE THEORY OF EVERYTHING

HERMAN ERIK

THE NOTORIOUS TRUE STORY OF THE KRAY TWINS VERLINDE





THE OFTHE THE MAKING OF THE ULTIMATE THEORY Congratulations & ADT

RULAN SIRI FY



Alexandre Belin, RCM, Shan-Ming Ruan, Gabor Sarosi & Antony Speranza [arXiv:2111.02429; arXiv:2208.0xxxx]

Complexity = Anything

Alexandre Belin, RCM, Shan-Ming Ruan, Gabor Sarosi & Antony Speranza [arXiv:2111.02429; arXiv:2208.0xxxx]

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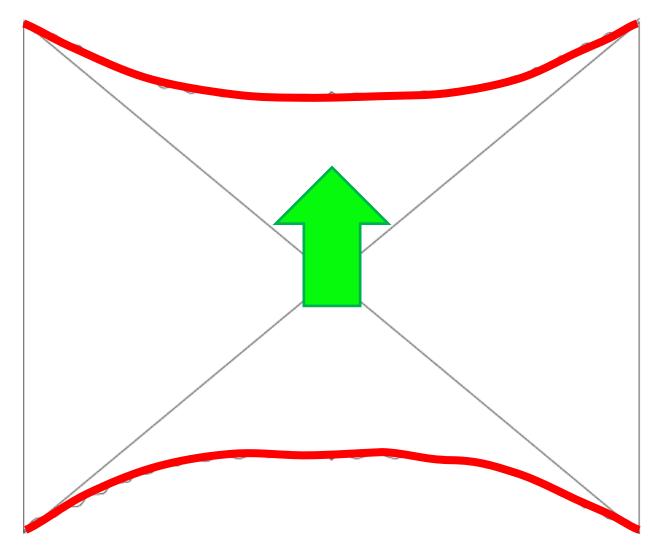
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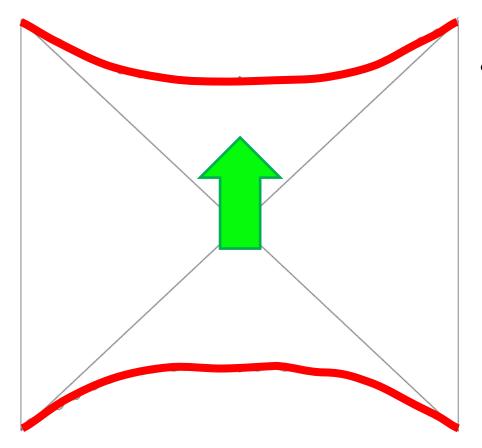
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• "to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity."



Entropy Susskind: Entanglement Ais not enough!

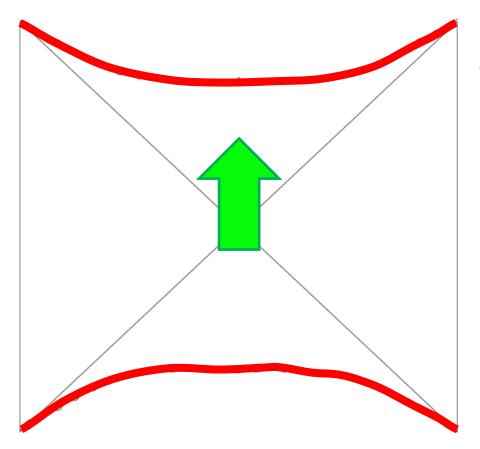
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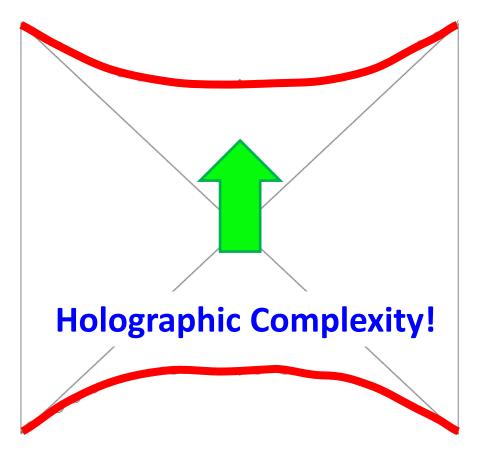
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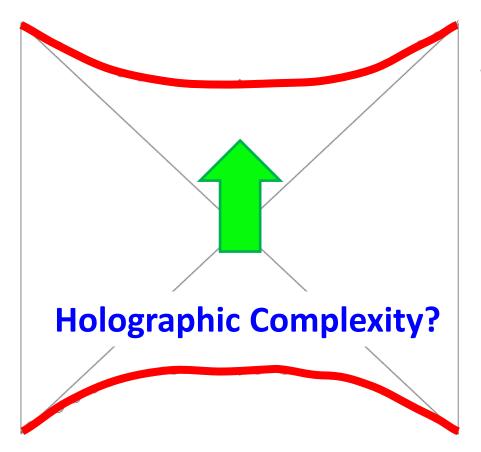
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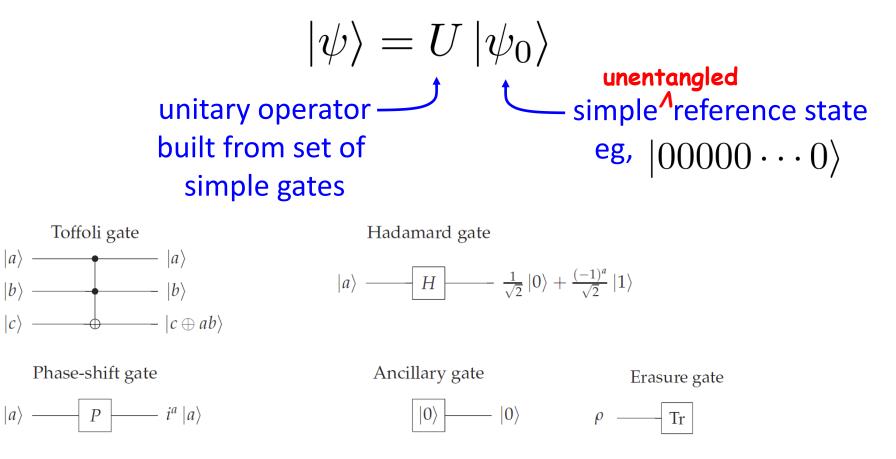
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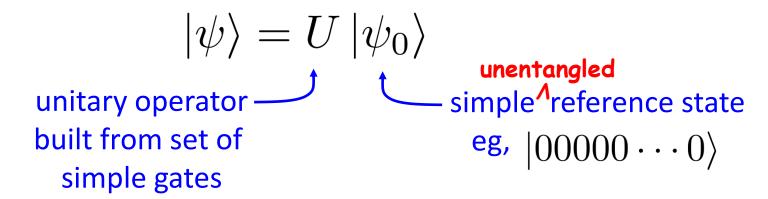
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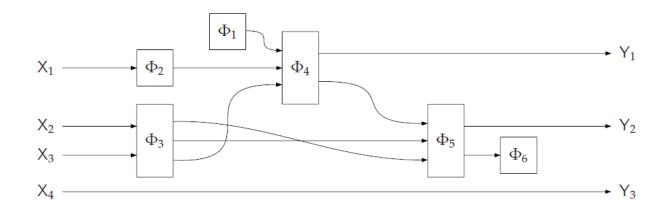
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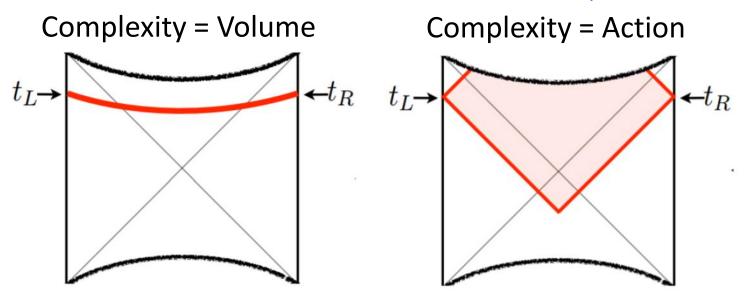
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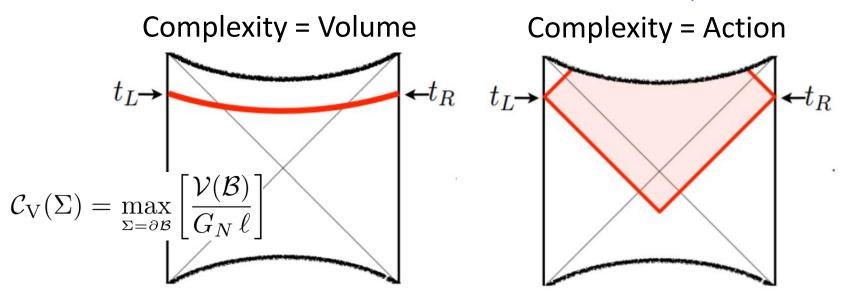
go beyond questions like `poly vs exp'

 <u>complexity=volume</u>: evaluate proper volume of extremal codim-one surface connecting Cauchy surfaces in boundary theory (cf holo EE) (Stanford & Susskind)



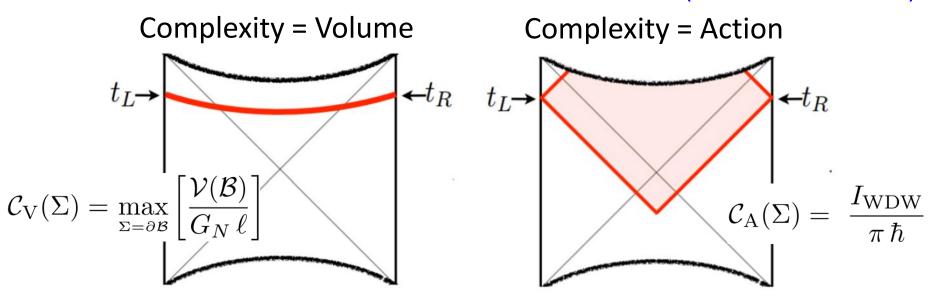
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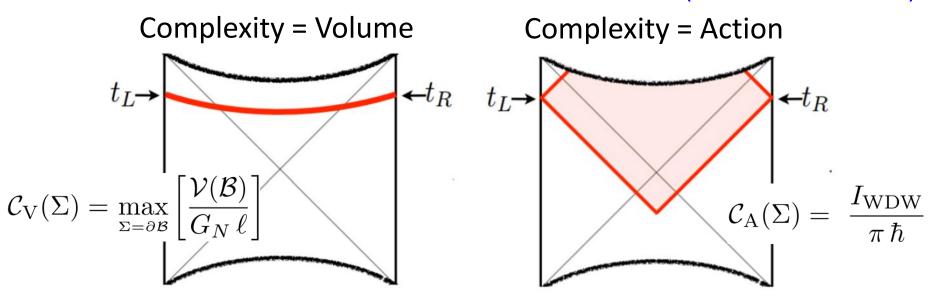
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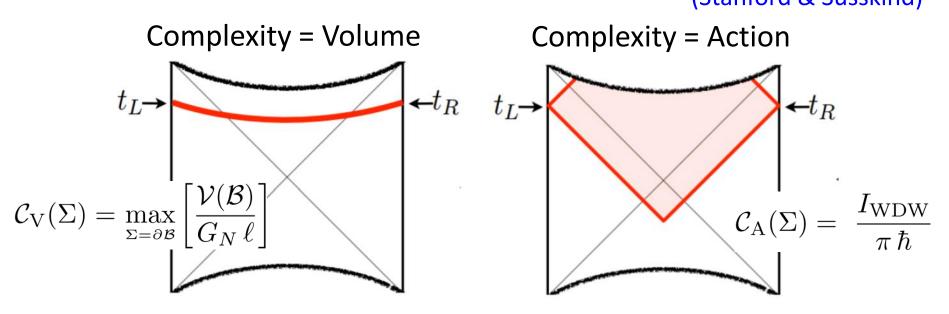
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- both of these gravitational "observables" probe the black hole interior (at arbitrarily late times on boundary)

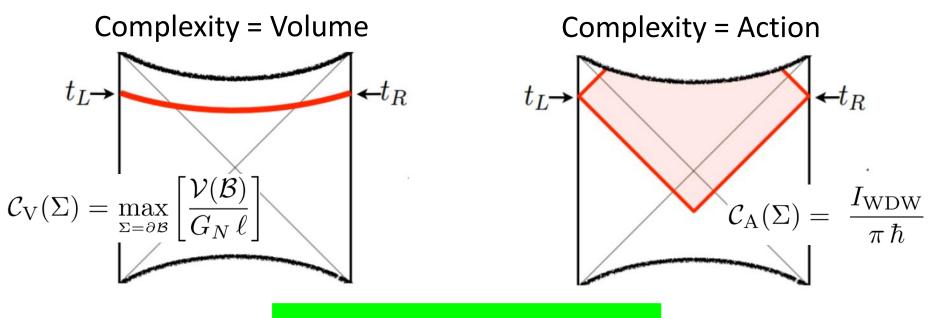
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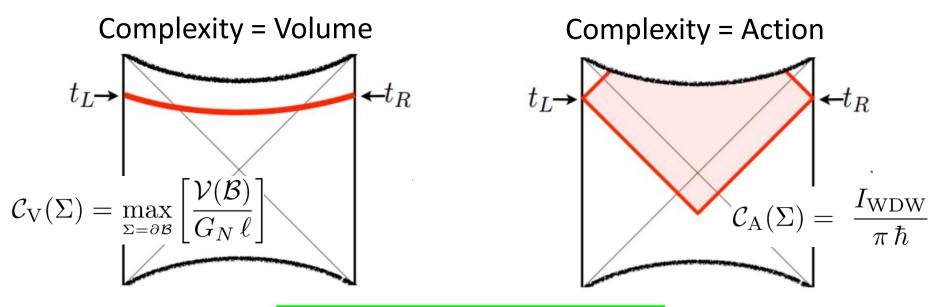
• <u>complexity=volume2.0</u>: evaluate spacetime volume of WDW patch $C'_V(\Sigma) = \frac{V_{WDW}}{G_N \ell^2}$ (Couch, Fischler & Nguyen)

Holographic Complexity:



WHY COMPLEXITY??

Holographic Complexity:



WHY COMPLEXITY??

- connection of complexity=volume to AdS/MERA
- linear growth (at late times)

$$\left. rac{d \mathcal{C}_{\mathrm{V}}}{dt}
ight|_{t
ightarrow \infty} = rac{8 \pi}{d-1} \, M$$
 (planar)

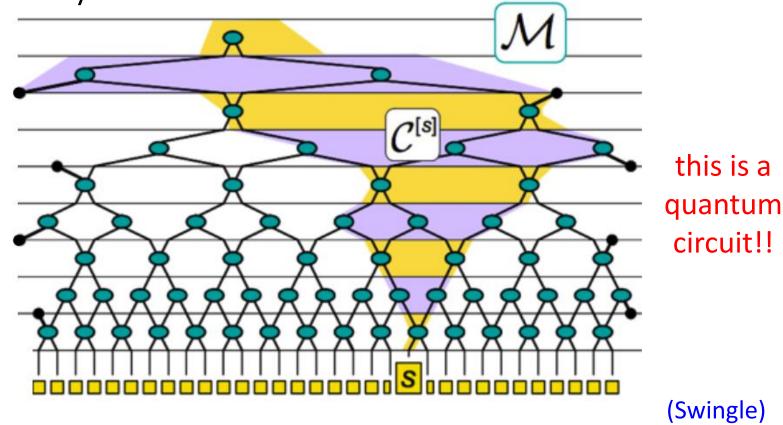
(d = boundary dimension)

$$\frac{d\mathcal{C}_{\mathcal{A}}}{dt}\Big|_{t\to\infty} = \frac{2M}{\pi}$$

"switchback" effect (information scrambling)

AdS/MERA:

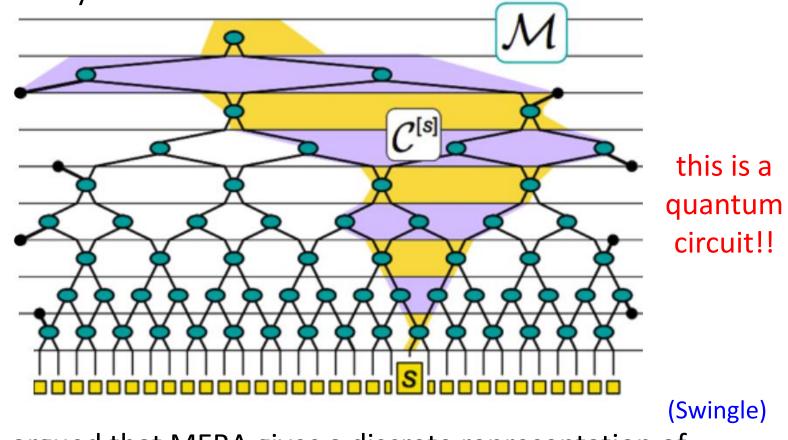
• MERA (Multi-scale Entanglement Renormalization Ansatz) provides efficient tensor network representation of ground-state wave-function in d=2 critical systems (Vidal; Vidal & Evenbly)



 has been argued that MERA gives a discrete representation of a time slice in AdS space!

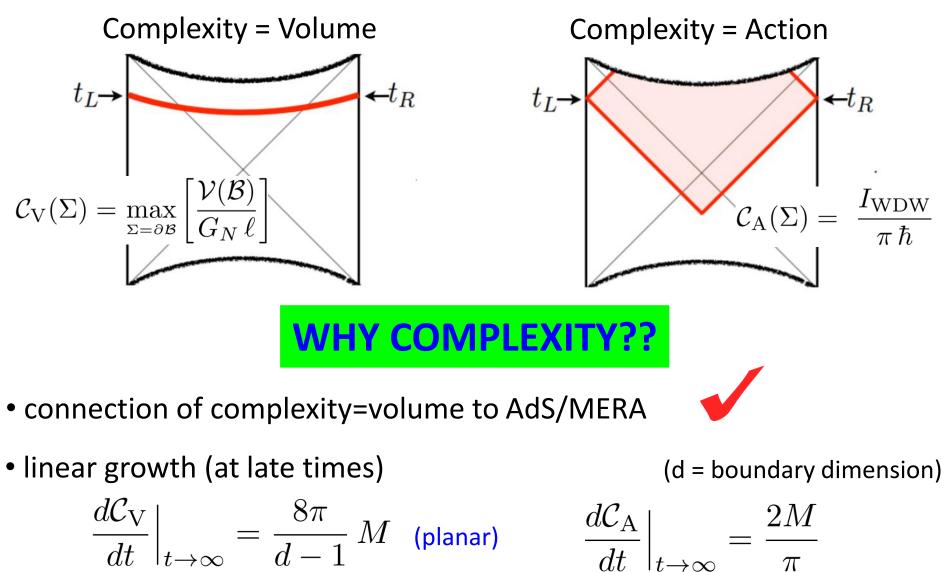
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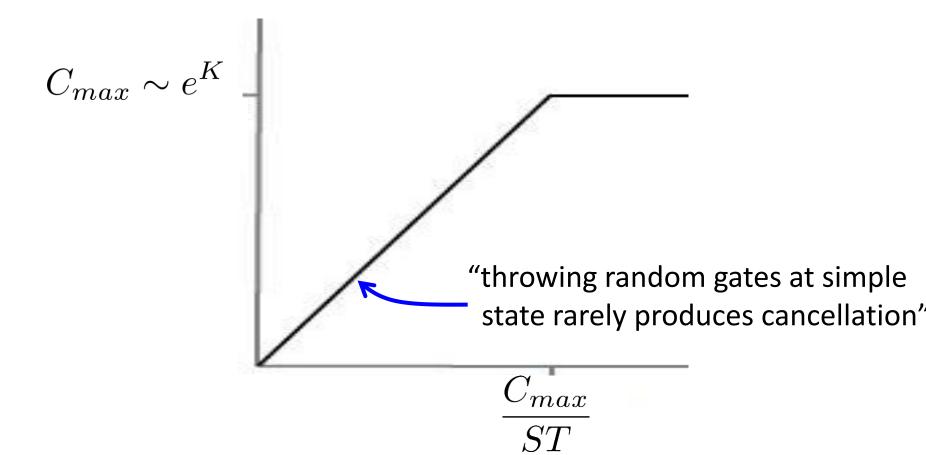
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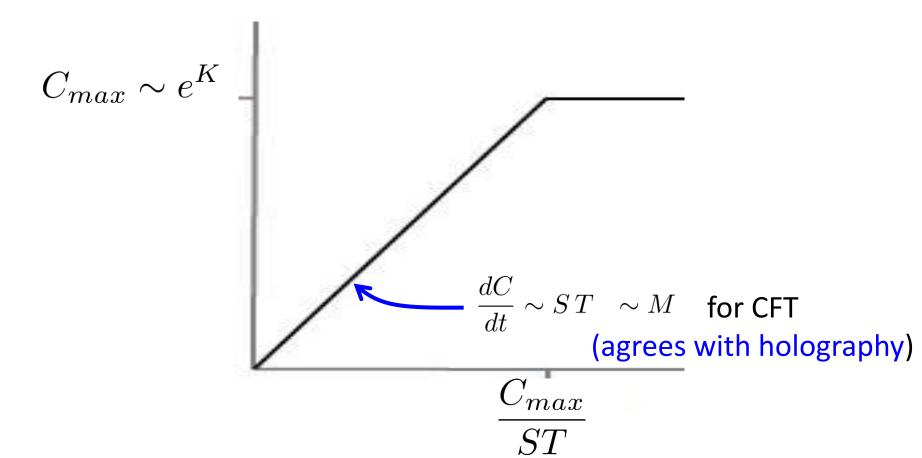
Linear Growth (at late times):

- large number of degrees of freedom + Hamiltonian is chaotic
 for very long time, growth is linear in time
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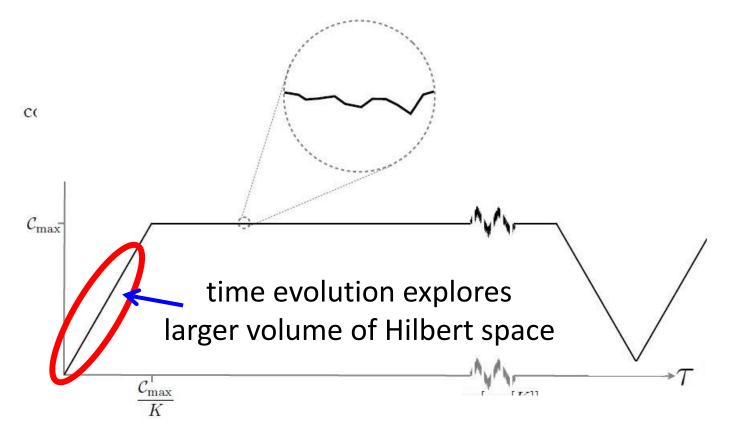
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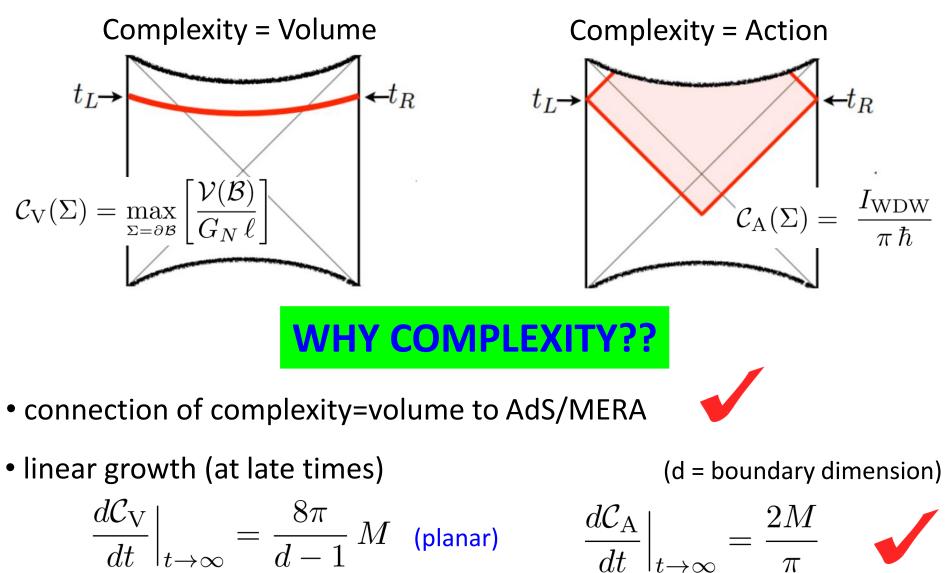
Susskind, Brown, . . .

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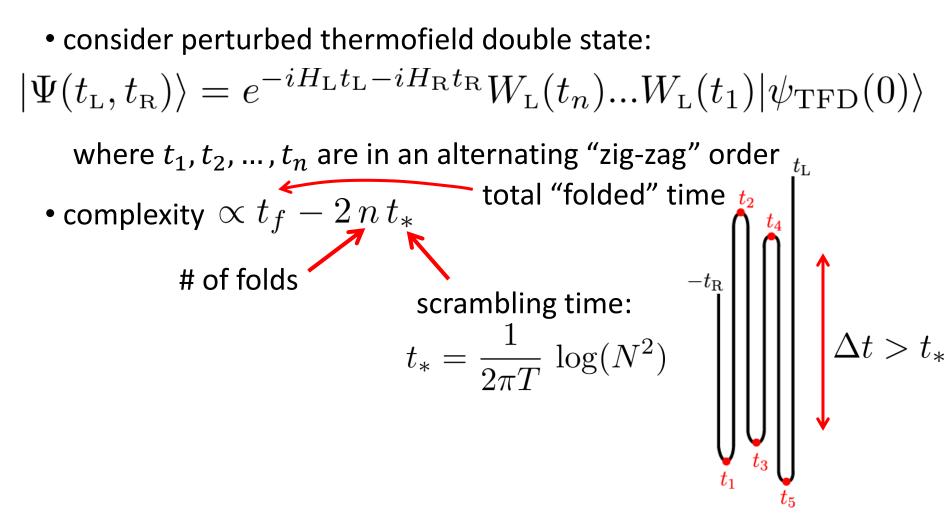
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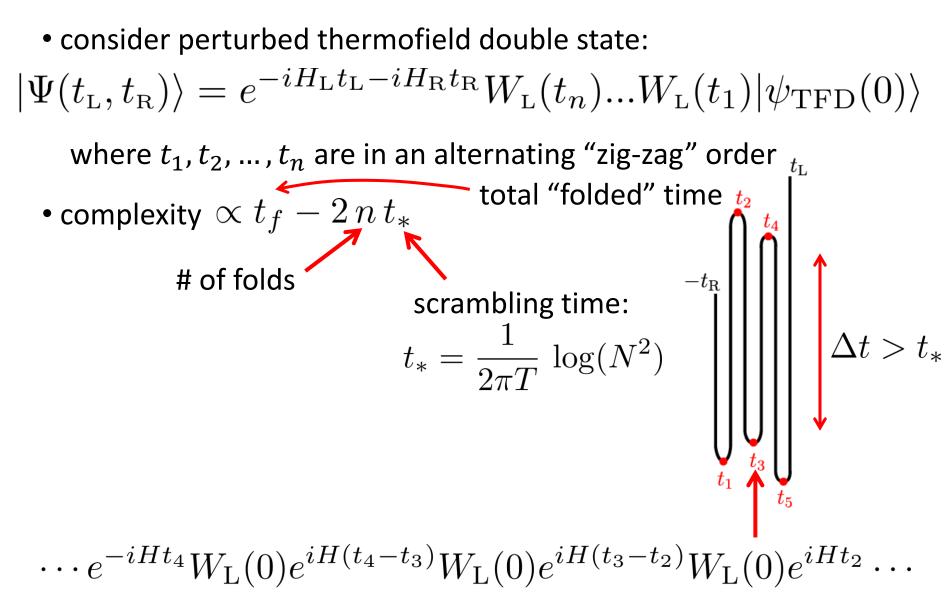
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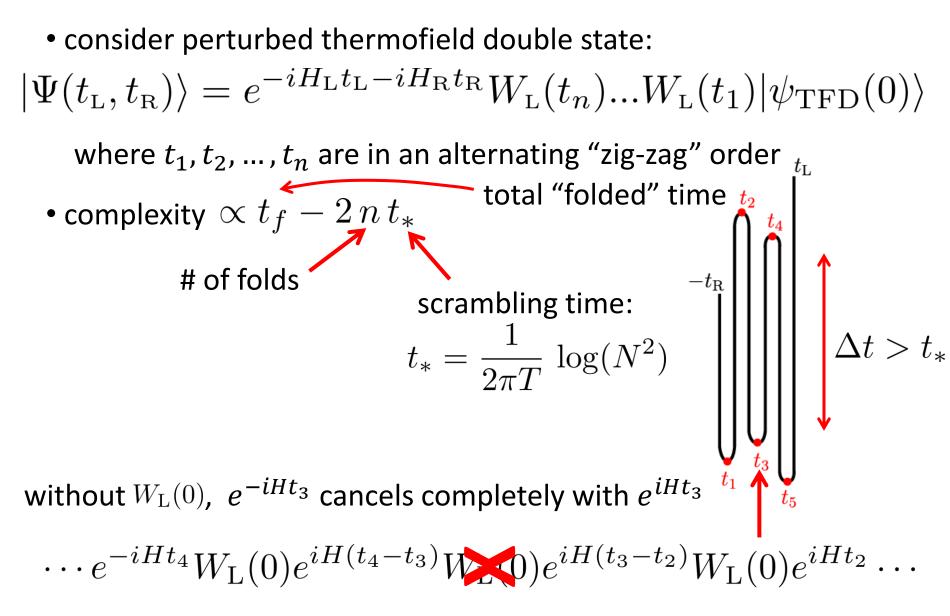
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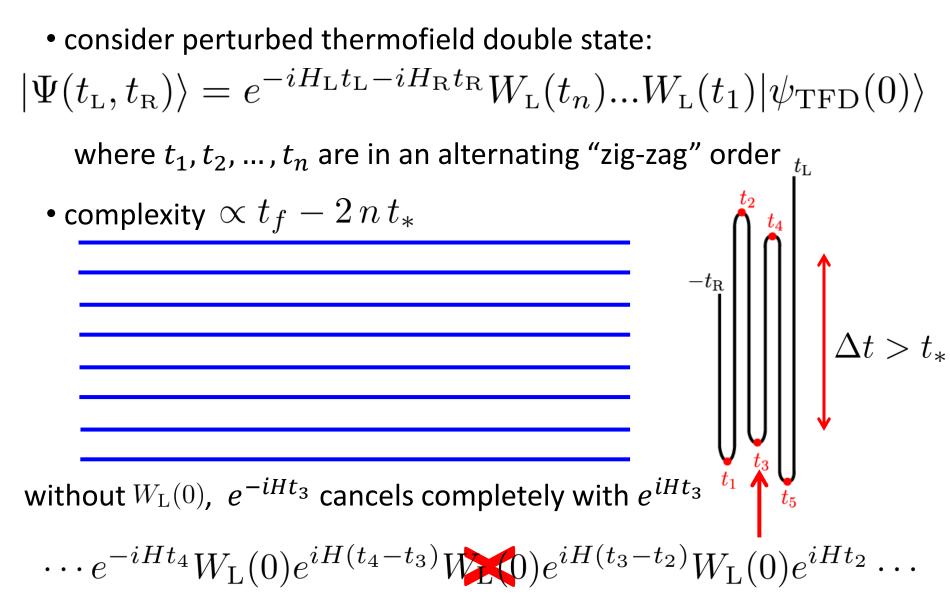
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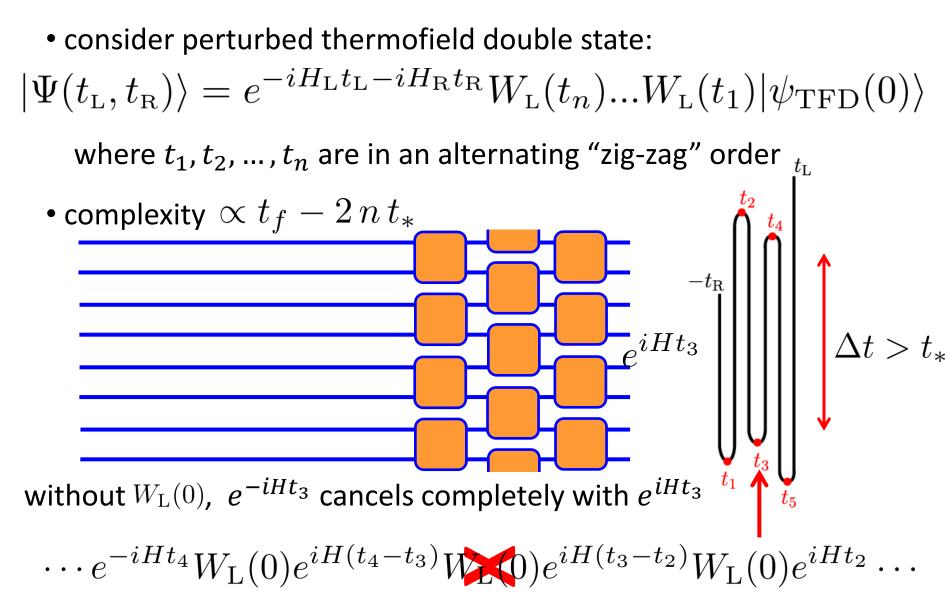
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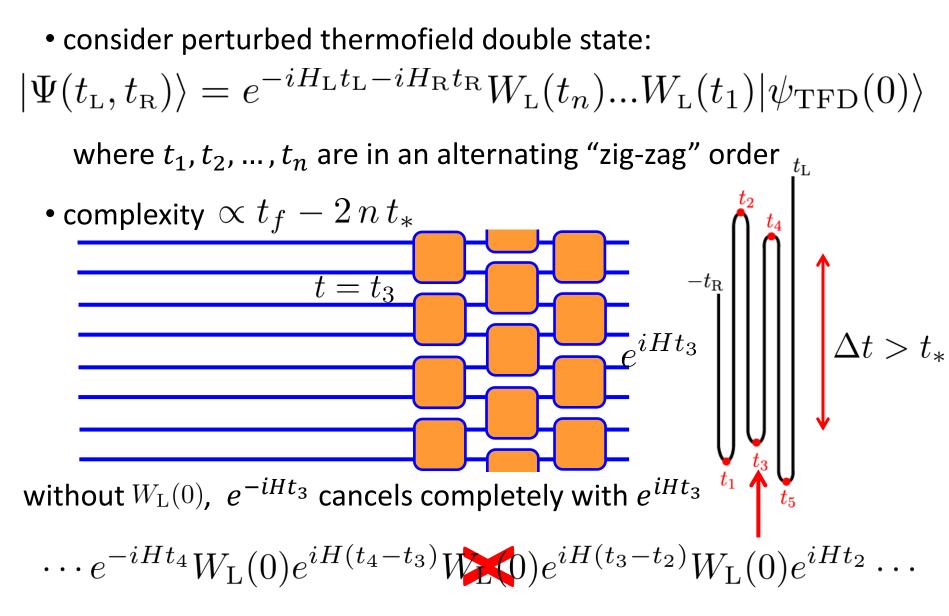


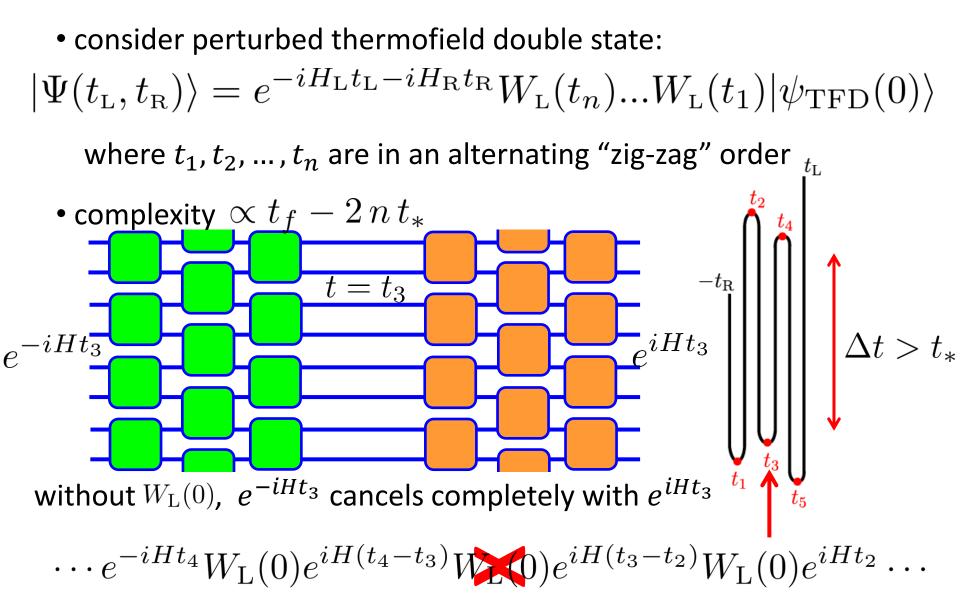


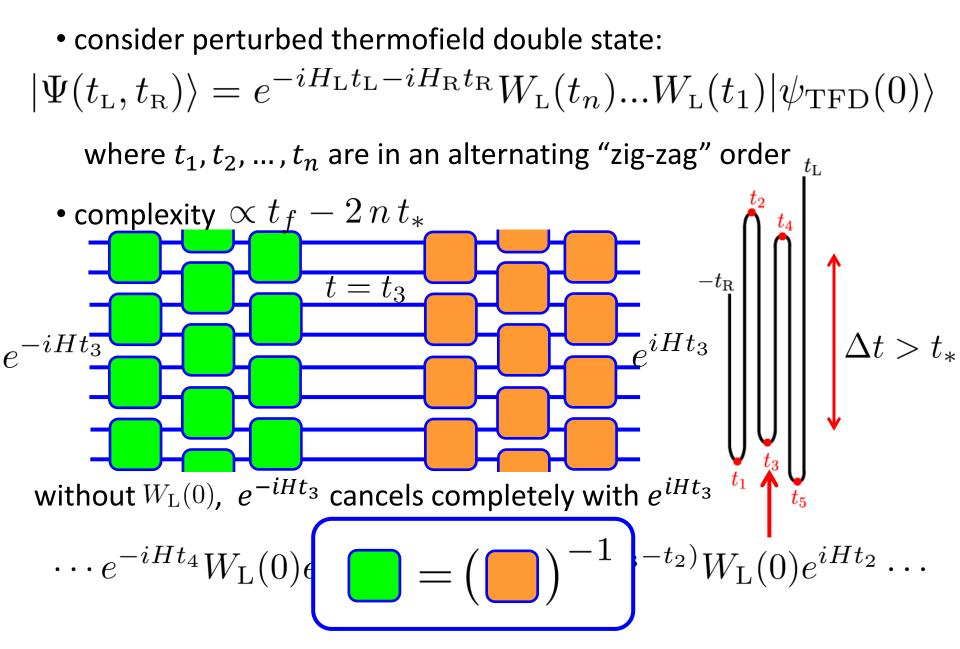


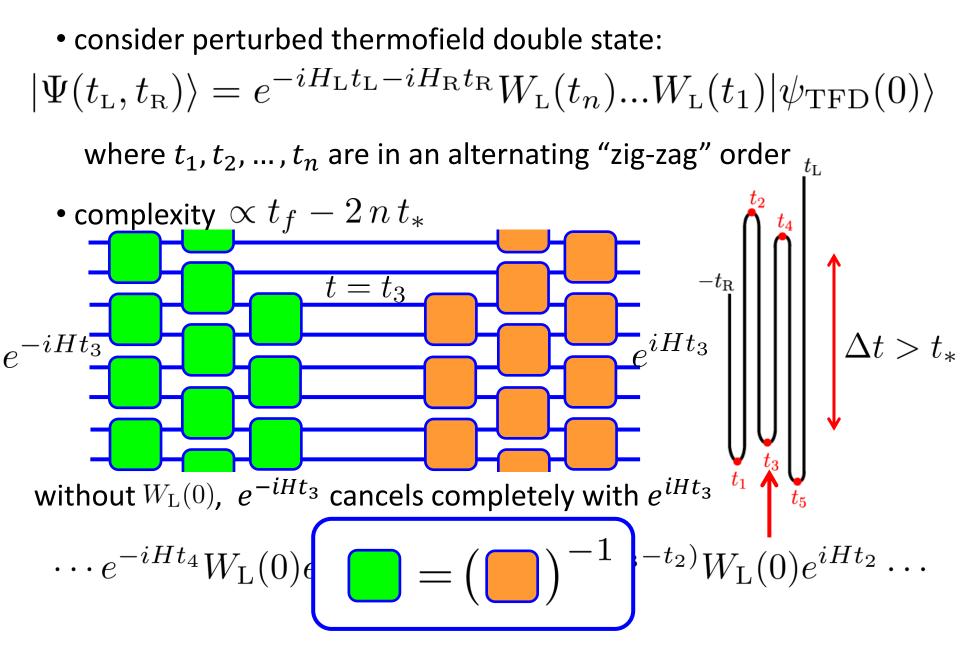


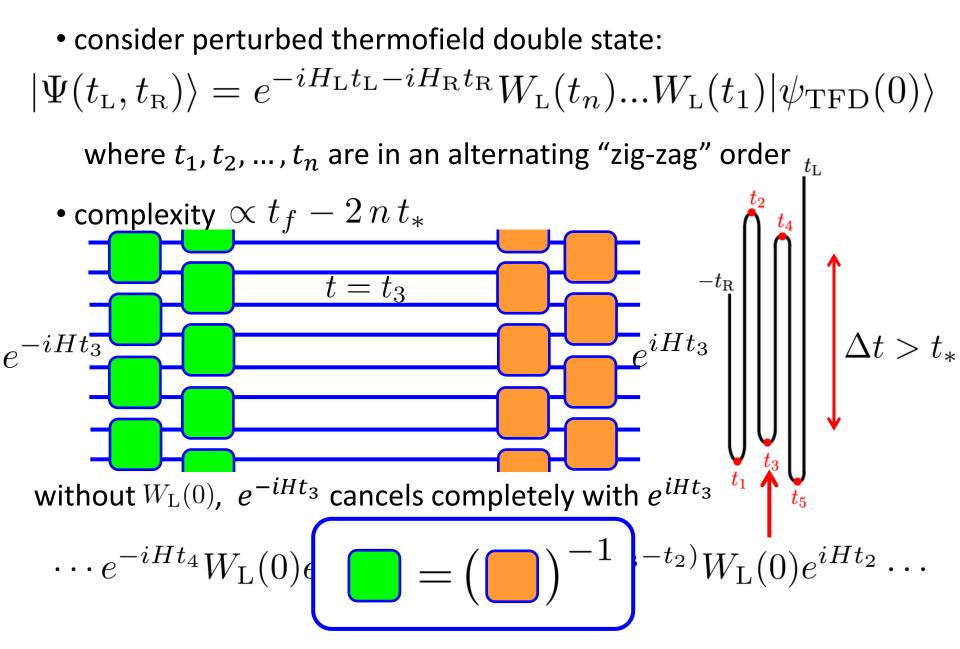


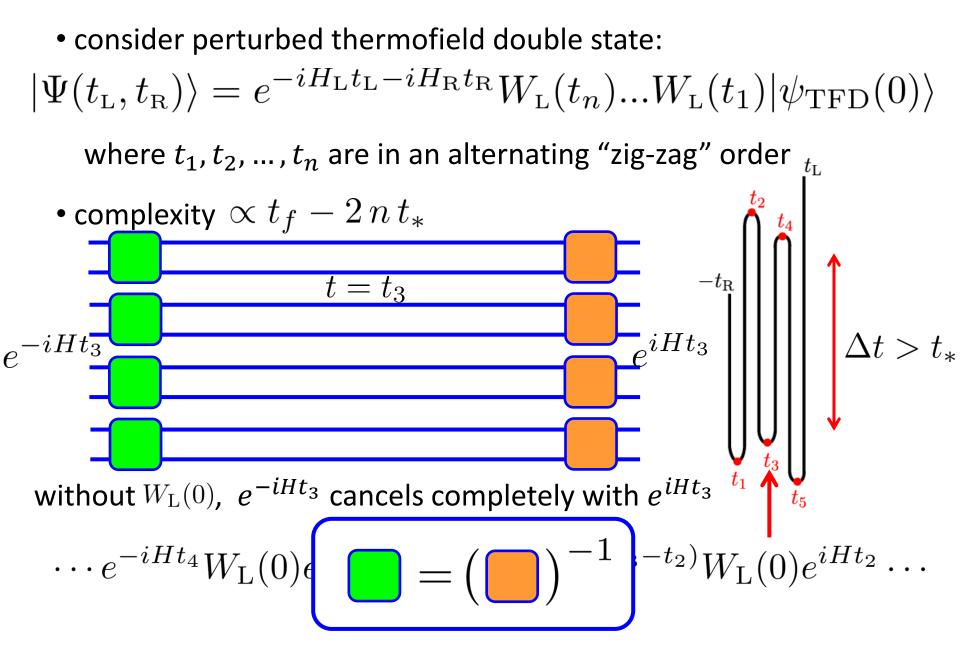


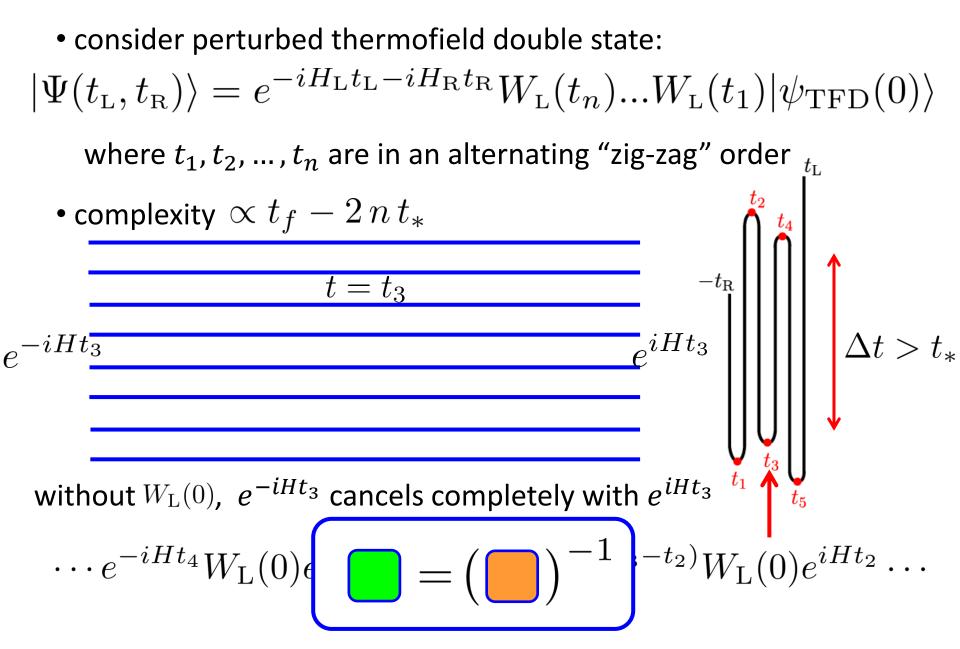






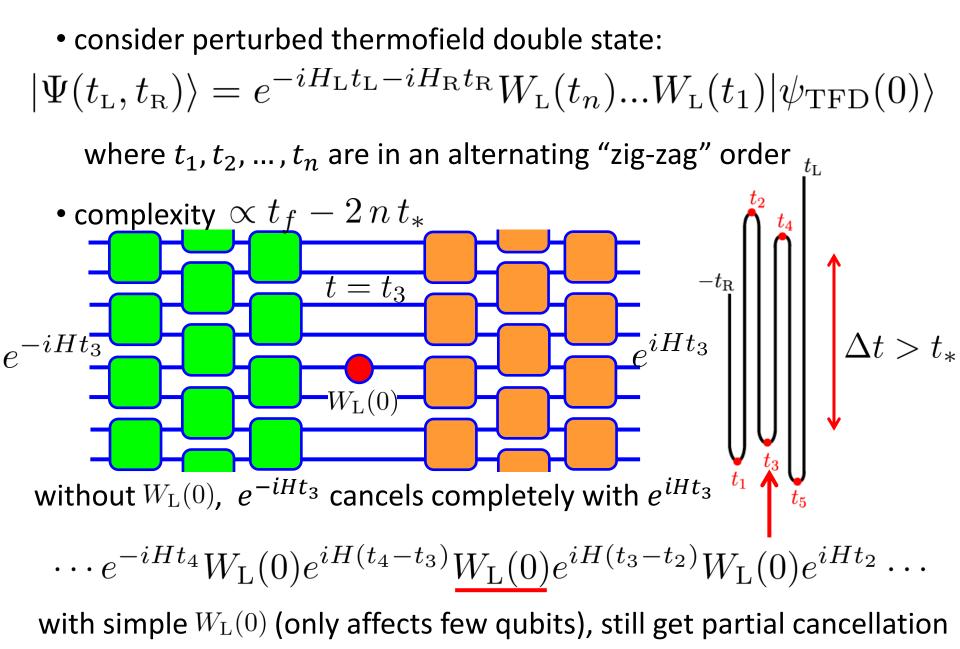


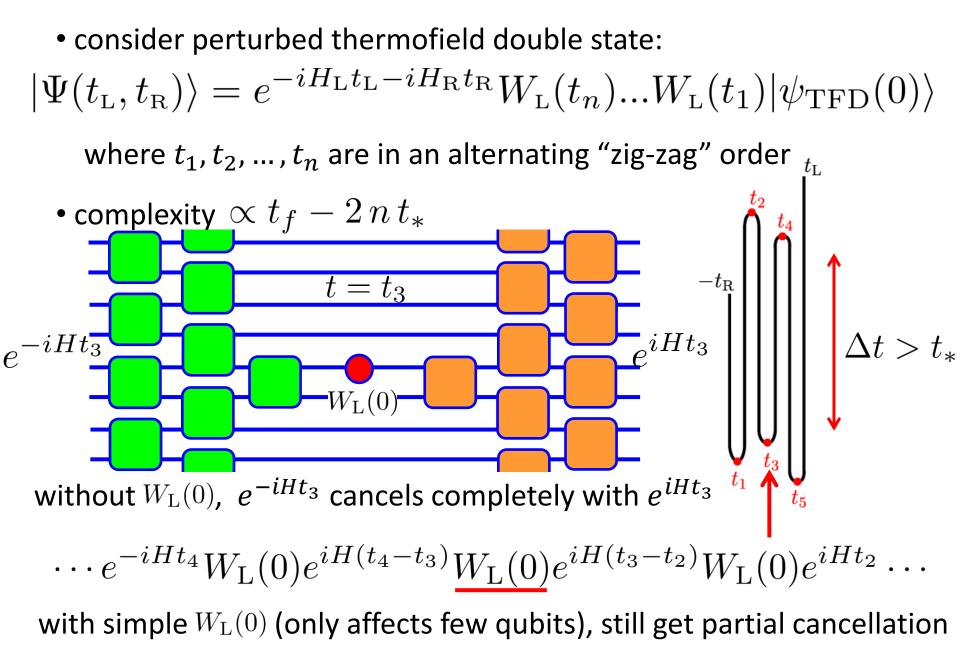


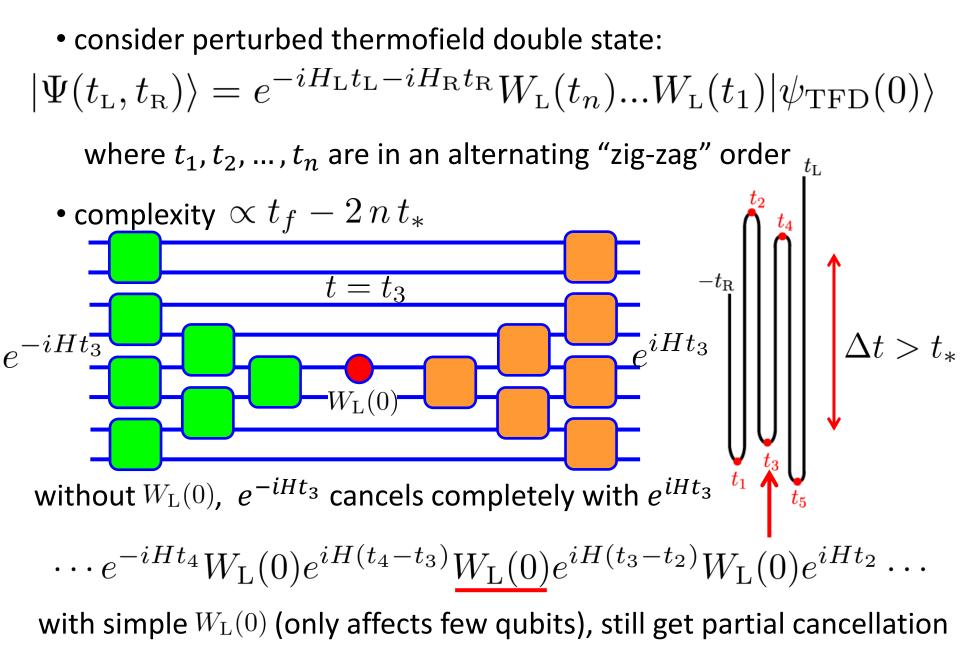


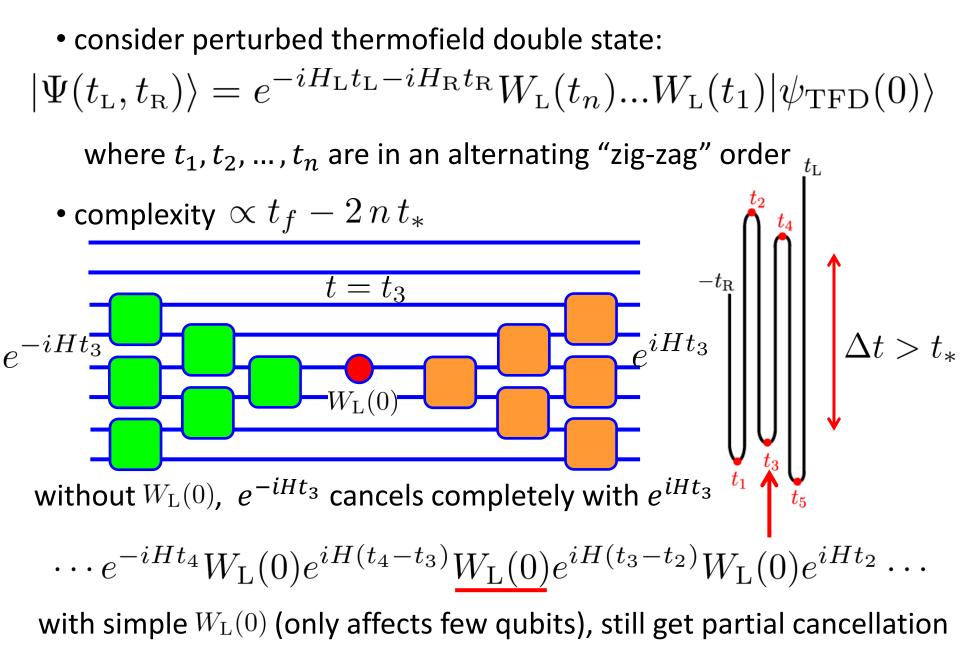
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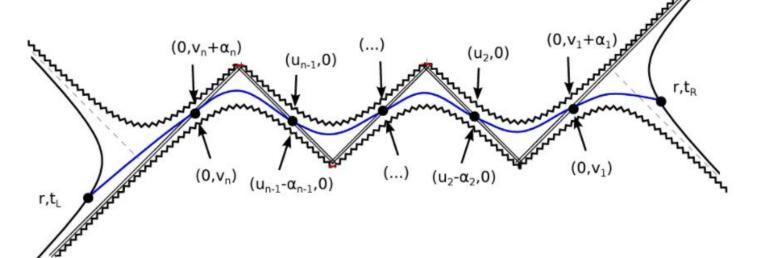
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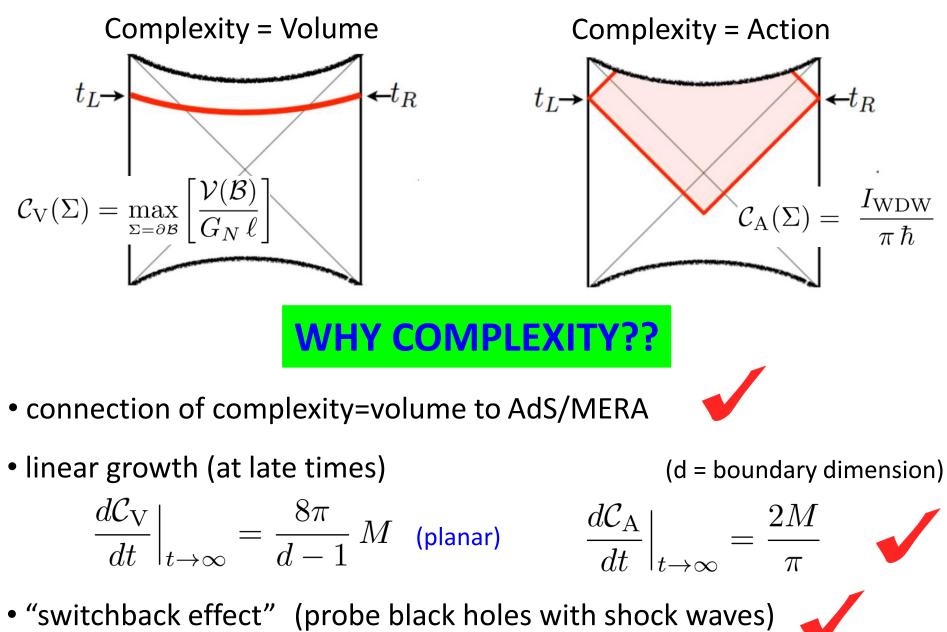
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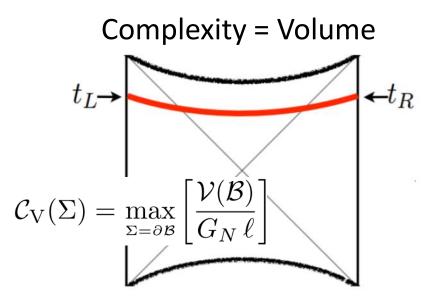
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- reproduce this behaviour by probing black hole with shock waves

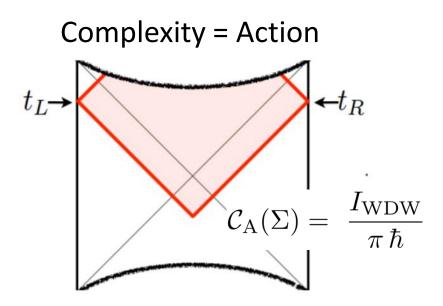


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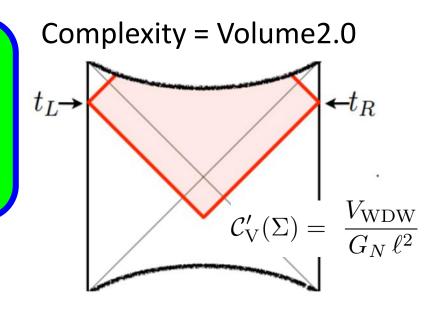


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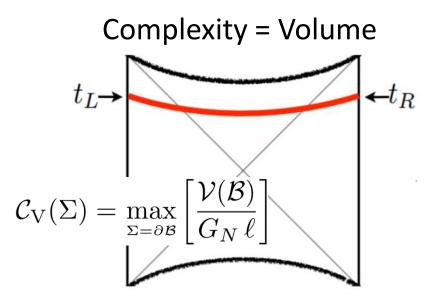


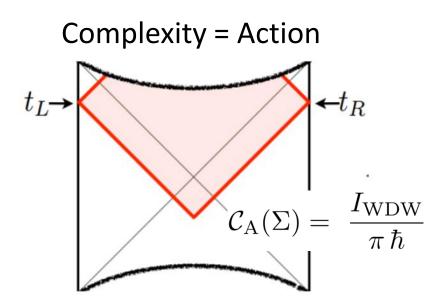


WHY Volume or Action or Spacetime Volume???



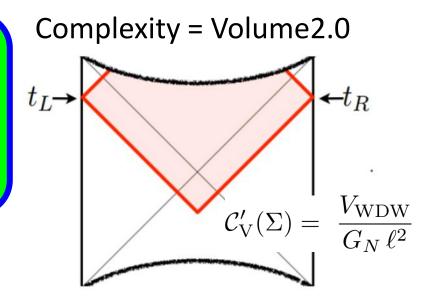
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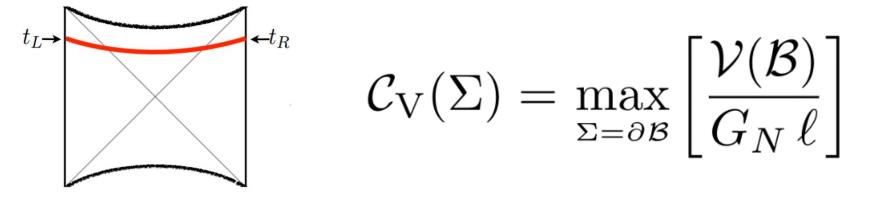
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Ambiguities in defining complexity?



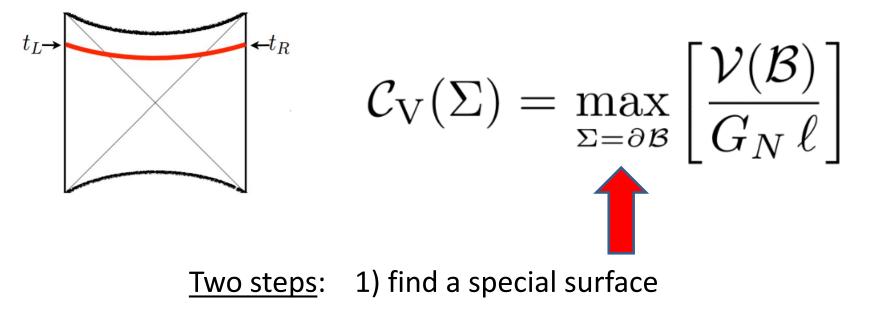
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 <u>complexity=volume</u>: evaluate proper volume of extremal codim-one surface connecting Cauchy surfaces in boundary theory (cf holo EE) (Stanford & Susskind)



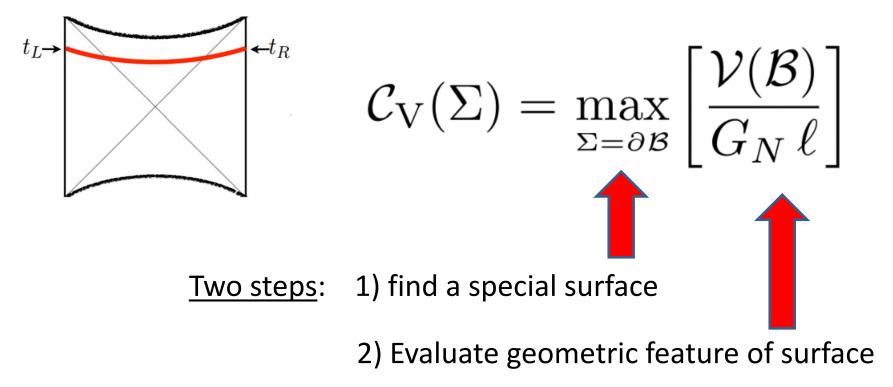
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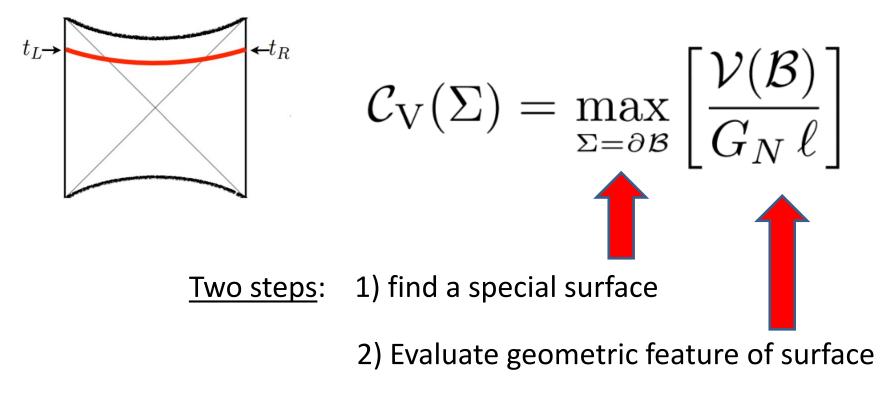
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• yields "nice" diffeomorphism invariant observable

 $t_{I} \rightarrow$

 $\leftarrow t_R$

Generalize two step procedure:

1) find a special surface Σ :

$$\delta_X \left(\int_{\Sigma} d^d \sigma \sqrt{h} \ F_2(g_{\mu\nu}; X^{\mu}) \right) = 0$$

• F_2 is scalar function of bkgd metric $g_{\mu\nu}$ and embedding $X^{\mu}(\sigma)$

 $t_L \rightarrow$

 $\leftarrow t_R$

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2) evaluate geometric feature of surface:

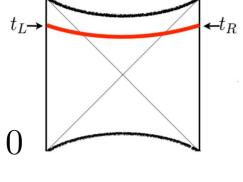
$$O_{F_1, \Sigma_{F_2}}(\Sigma_{CFT}) = \frac{1}{G_N L} \int_{\Sigma_{F_2}} d^d \sigma \sqrt{h} \ F_1(g_{\mu\nu}; X^{\mu})$$

• F_1 is *scalar* function of bkgd metric $g_{\mu\nu}$ and embedding $X^{\mu}(\sigma)$

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$$O_{F_1,\Sigma_{F_2}}(\Sigma_{CFT}) = \frac{1}{G_N L} \int_{\Sigma_{F_2}} d^d \sigma \sqrt{h} F_1(g_{\mu\nu}; X^{\mu})$$

- F_1 is *scalar* function of bkgd metric $g_{\mu\nu}$ and embedding $X^{\mu}(\sigma)$
- yields "nice" diffeomorphism invariant observable

 $t_L \rightarrow$

 $\leftarrow t_R$

Generalize two step procedure:

1) find a special surface Σ :

$$\delta_X \left(\int_{\Sigma} d^d \sigma \sqrt{h} \ F_2(g_{\mu\nu}; X^{\mu}) \right) = 0$$



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So what?

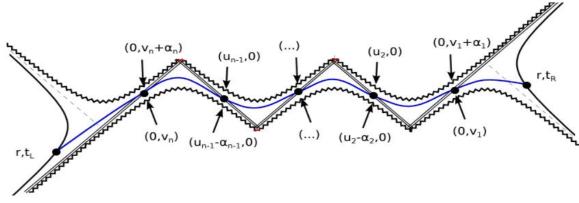
So what?

1) Observables grow linearly with time at late times:

$$\lim_{\tau \to \infty} O_{F_1, \Sigma_{F_2}}(\tau) \sim P_{\infty} \tau$$

where in large *T* limit, the constant $P_{\infty} \propto \text{mass}$

2) Observables exhibit "switchback effect", ie, universal time delay in response to shock waves falling into the dual black hole



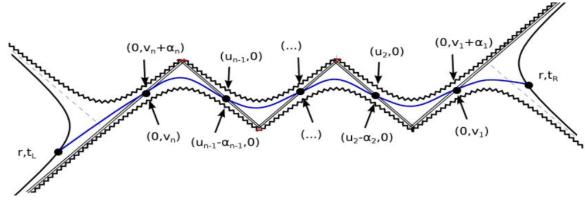
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 when F₁ and F₂ constructed with bkgd curvtures, analysis similar to extremal volume **Simple Example:** $F_1 = F_2 = 1 + \lambda L^4 C_{abcd} C^{abcd}$

profile determined by classical mechanics problem

 $\dot{r}^2 + \widetilde{U}(r) = P_v^2$ with $\widetilde{U}(r) = -f(r)a^2(r)\left(\frac{r}{\tau}\right)^{2(d-1)}$, where $f(r) = \frac{r^2}{L^2} \left(1 - \frac{r_h^d}{r^d} \right)$ and $a(r) = 1 + \tilde{\lambda} \left(\frac{r_h}{r} \right)^{2d}$ $\tilde{U}(r)$ $P_v^2 \ge P_\infty^2$ P^2_{∞} $\cdots P_{\nu}^2(t'_{\rm R})$ • turning point: $r_{\rm min}$ $P_{\nu}^2(t_{\rm R})$ $P_v^2 = \widetilde{U}(r_{min})$ r_{min} $P_{\nu}^2 \leq P_{\infty}^2$ $P_{\infty}^{2} = U(r_{f})$ $- = \xi \left(\frac{r_{h}}{I}\right)^{2d}$ () $\widetilde{U}'(r_f) = 0$

So what?

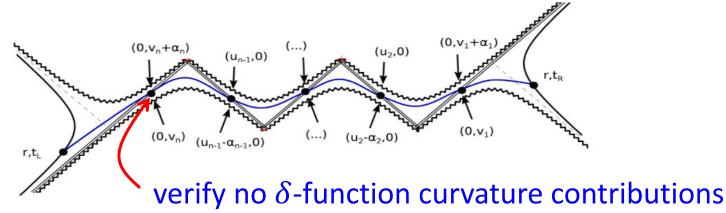
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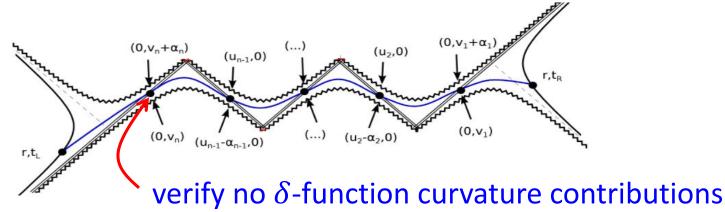
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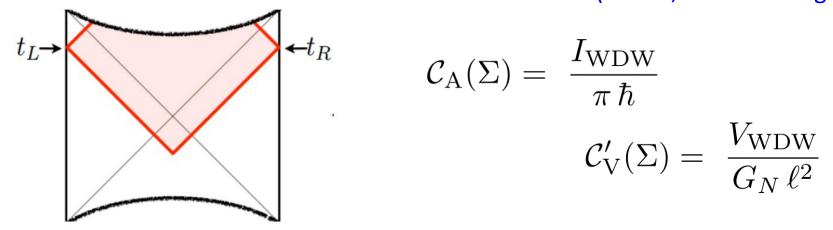
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What about extensions of CA and CV2.0?

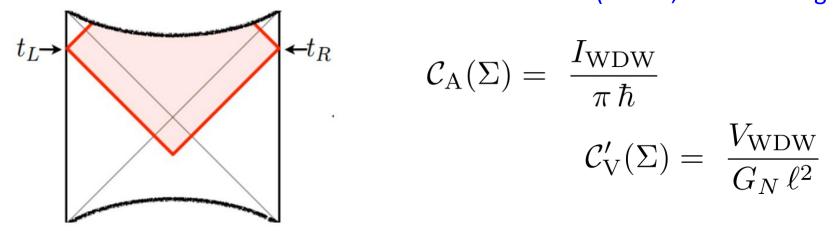
- <u>complexity=action</u>: evaluate gravitational action for Wheeler-DeWitt patch = domain of dependence of bulk time slice connecting boundary Cauchy slices in CFT (Brown, Roberts, Swingle, Susskind & Zhao)
- <u>complexity=volume2.0</u>: evaluate spacetime volume of WDW patch (Couch, Fischler & Nguyen)



• no extremization procedure implemented; surfaces bounding volume (ie, codimension-zero region) are light sheets

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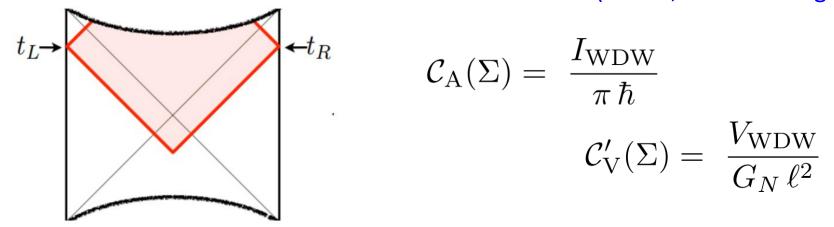
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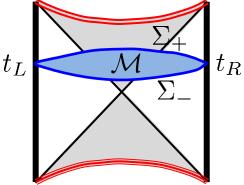


<u>Two steps</u>: 1) find a special surfaces bounding codim.-0 region
2) Evaluate geometric feature of codim.-0 region
(& bounding surfaces)

• yields "nice" diffeomorphism invariant observable

Generalized procedure for codim.-0 observables:

1) find a bounding surfaces Σ_{\pm} :



$$\delta_{\{X_{+},X_{-}\}} \Big(\int_{\Sigma_{+}} d^{d}\sigma \sqrt{h} F_{4}(g_{\mu\nu};X_{+}^{\mu}) + \int_{\Sigma_{-}} d^{d}\sigma \sqrt{h} F_{5}(g_{\mu\nu};X_{-}^{\mu}) + \int_{\mathcal{M}} d^{d+1}\sigma \sqrt{-g} F_{6}(g_{\mu\nu}) \Big) = 0$$

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2) Evaluate geometric feature of corresponding region:

$$O(\Sigma_{CFT}) = \frac{1}{G_N L} \int_{\Sigma_+} d^d \sigma \sqrt{h} F_1(g_{\mu\nu}; X^{\mu}_+)$$
$$-\frac{1}{G_N L} \int_{\Sigma_-} d^d \sigma \sqrt{h} F_2(g_{\mu\nu}; X^{\mu}_-) + \frac{1}{G_N L^2} \int_{\mathcal{M}} d^{d+1} \sigma \sqrt{-g} F_3(g_{\mu\nu})$$

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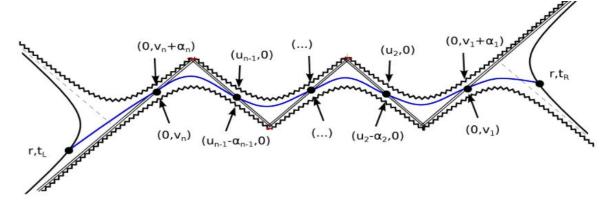
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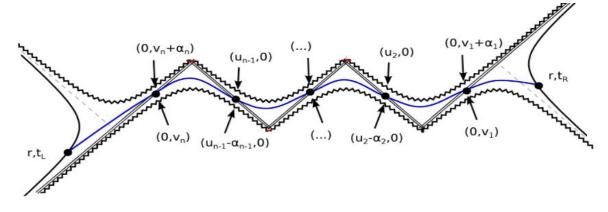
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Simplest Example:

extremize the functional

$$O(\Sigma_{CFT}) = \frac{\alpha_+}{G_N L} \int_{\Sigma_+} d^d \sigma \sqrt{h} + \frac{\alpha_-}{G_N L} \int_{\Sigma_-} d^d \sigma \sqrt{h} + \frac{1}{G_N L^2} \int_{\mathcal{M}} d^{d+1} \sigma \sqrt{-g}$$

 t_R

 t_L

- evaluating the volumes of the bounding surfaces Σ_{\pm} weighted by coefficients α_{\pm} , as well as of volume of codim.-0 region \mathcal{M}
- extremal equations yields CMC surfaces (eg, see Witten)

$$K(\Sigma_{+}) = \frac{1}{\alpha_{+} L} \qquad \qquad K(\Sigma_{-}) = -\frac{1}{\alpha_{-} L}$$

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$$K(\Sigma_{+}) = \frac{1}{\alpha_{+} L} \qquad \qquad K(\Sigma_{-}) = -\frac{1}{\alpha_{-} L}$$

- in limit $\alpha_{\pm} \rightarrow 0$, these surfaces become the future/past light sheets $\longrightarrow \mathcal{M}$ becomes WDW patch!
- evaluate volume (same functional) ----> CV2.0
- evaluate action (including bdy terms) ----> CA

Conclusions/Questions/Outlook:

- simple example but "classical mechanics" analysis readily extends to $F_1(g_{\mu\nu}, \mathcal{R}_{\mu\nu\rho\sigma}, \nabla_{\mu})$ and to observables where $F_1 \neq F_2$
- couplings for curvature invariants should not be too large
- similar behaviour appears to hold for functionals including dependence on extrinsic curvature
- infinite class of holographic observables equally viable candidates for gravitational dual of complexity!!
- can freedom in constructing gravitational observables be related to freedom in constructing complexity model in boundary QFT
- is there something that singles out maximal volume?
- what is role of extremal solutions which are not global maxima and probe very near to singularity?
- further investigation of codimension-zero observables
- add matter contributions to new observables (eg, CA proposal)

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Lots to explore!

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THE OFTHE THE MAKING OF THE ULTIMATE THEORY Congratulations & ADT

RDIAN SIRIEY

