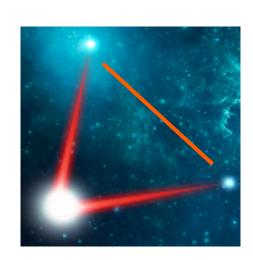
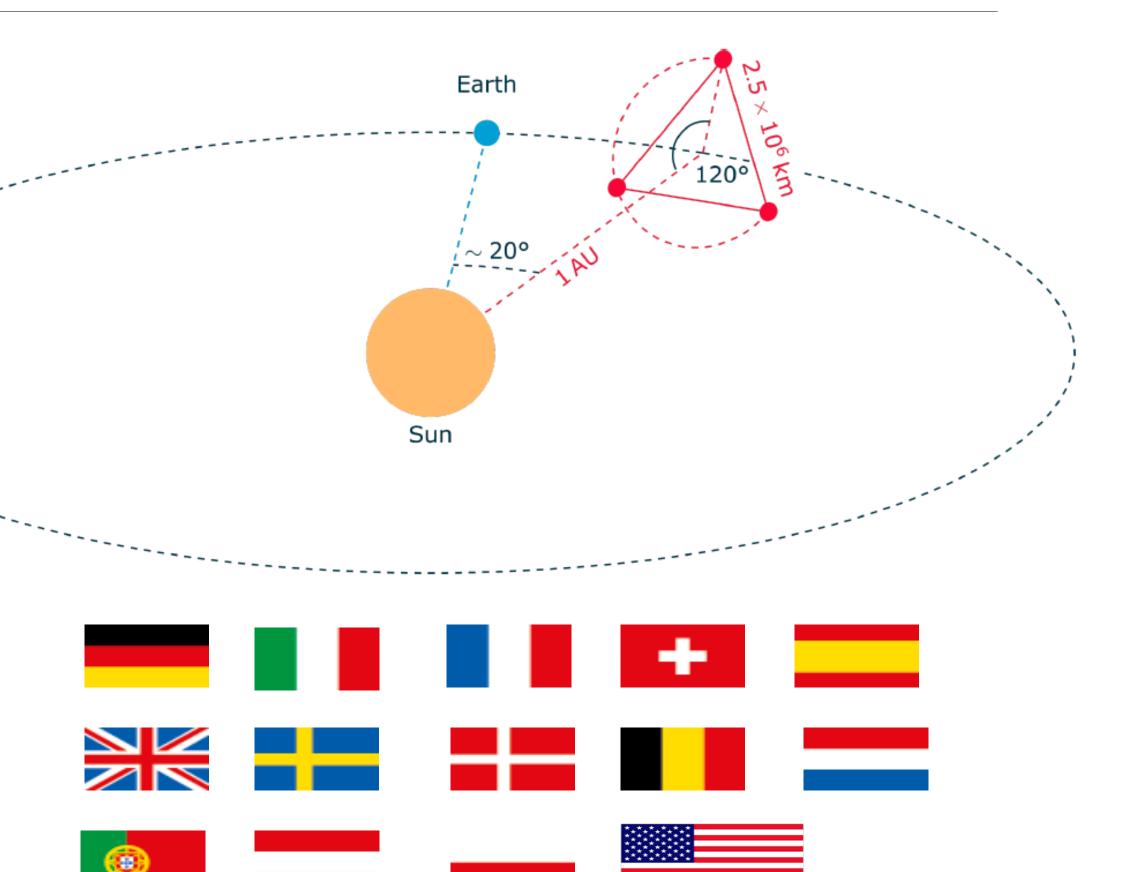
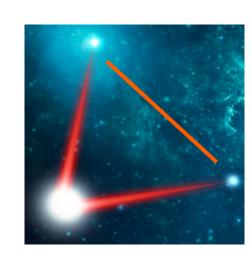
LASER INTERFEROMETER SPACE ANTENNA



The LISA mission

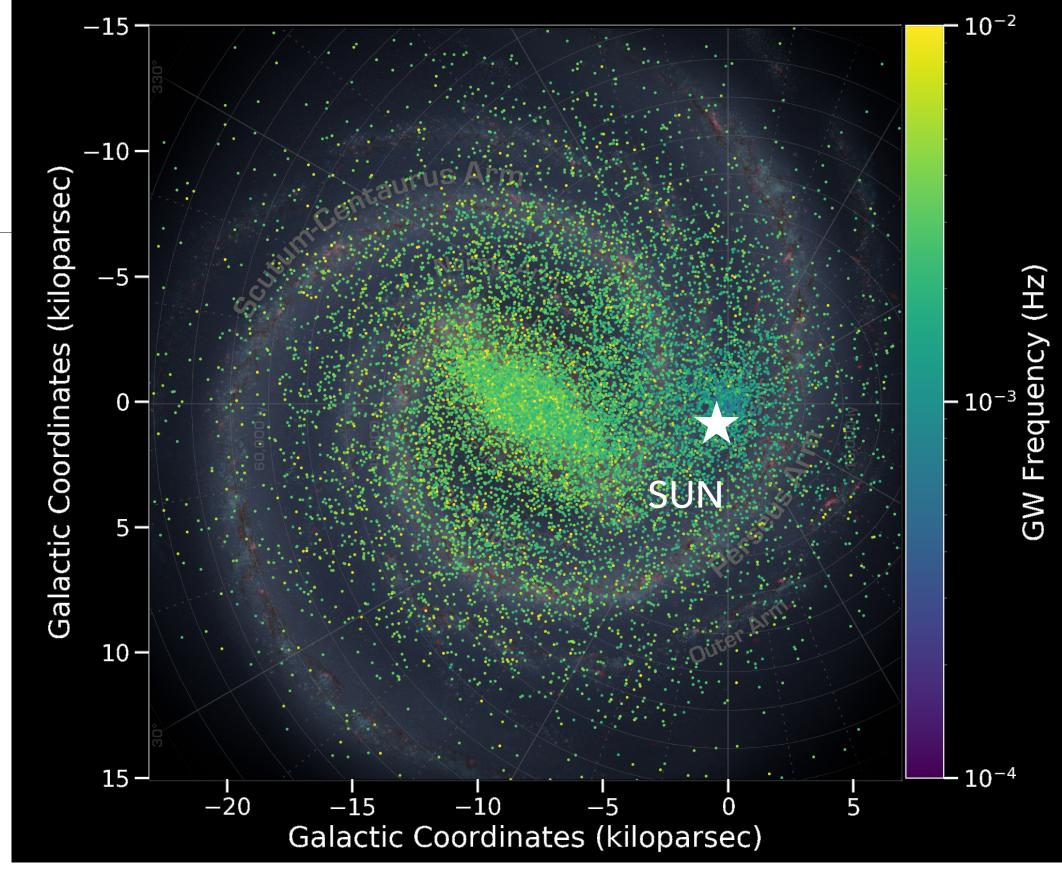
- Mission to detect Gravitational Waves with much lower frequency than LIGO/Virgo
- Three satellites in triangular formation, trailing Earth
- ESA lead, member state contributions, strong link with NASA
- Budget: ~2.6B€ (ESA: 1.5B€, NASA: ~600M€,Member states: ~500M€)
- Sensitive in mHz regime
- Launch in mid-2030s

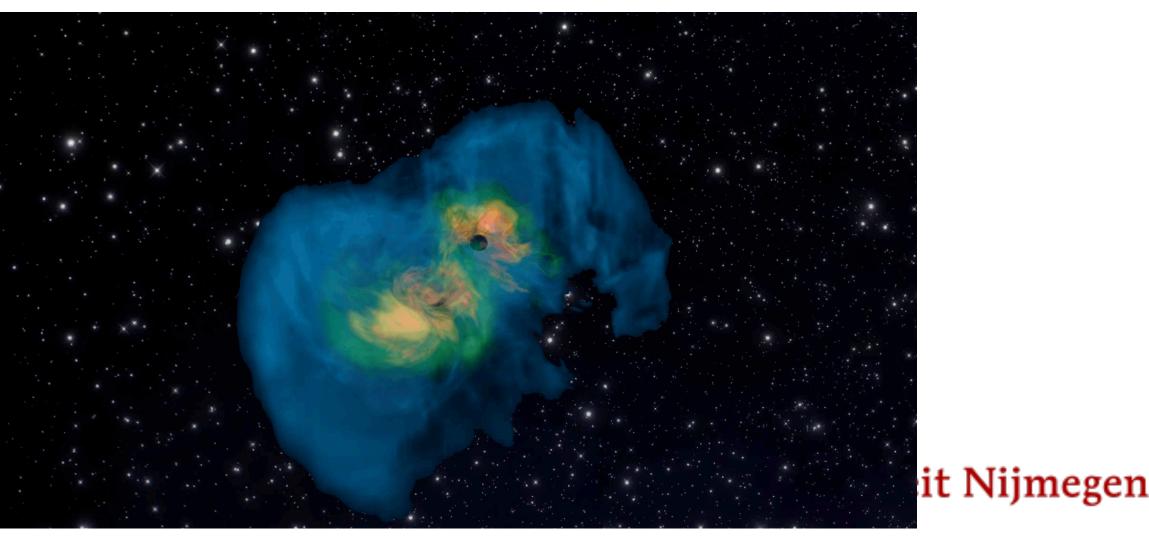


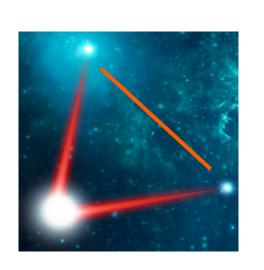


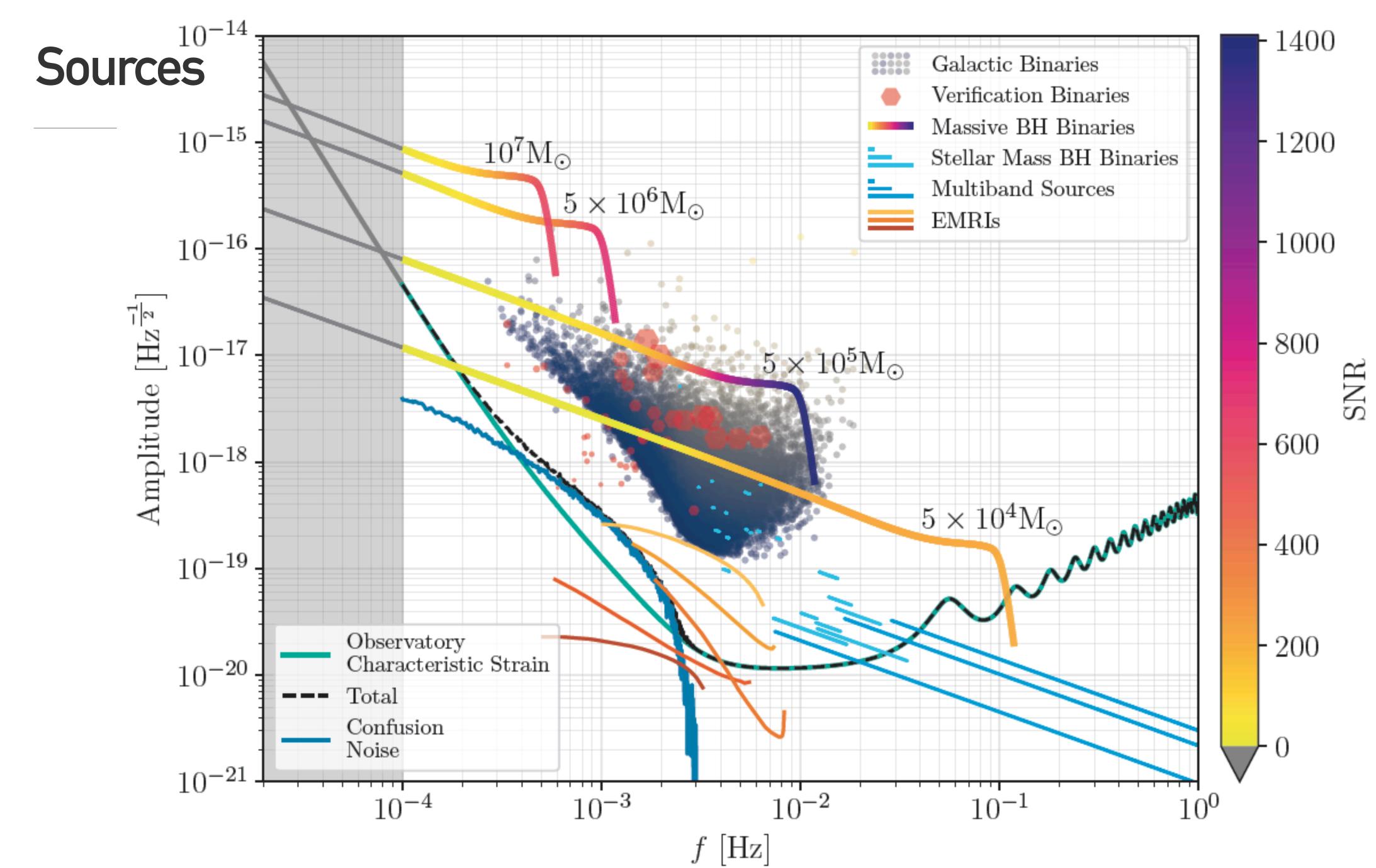
LISA science

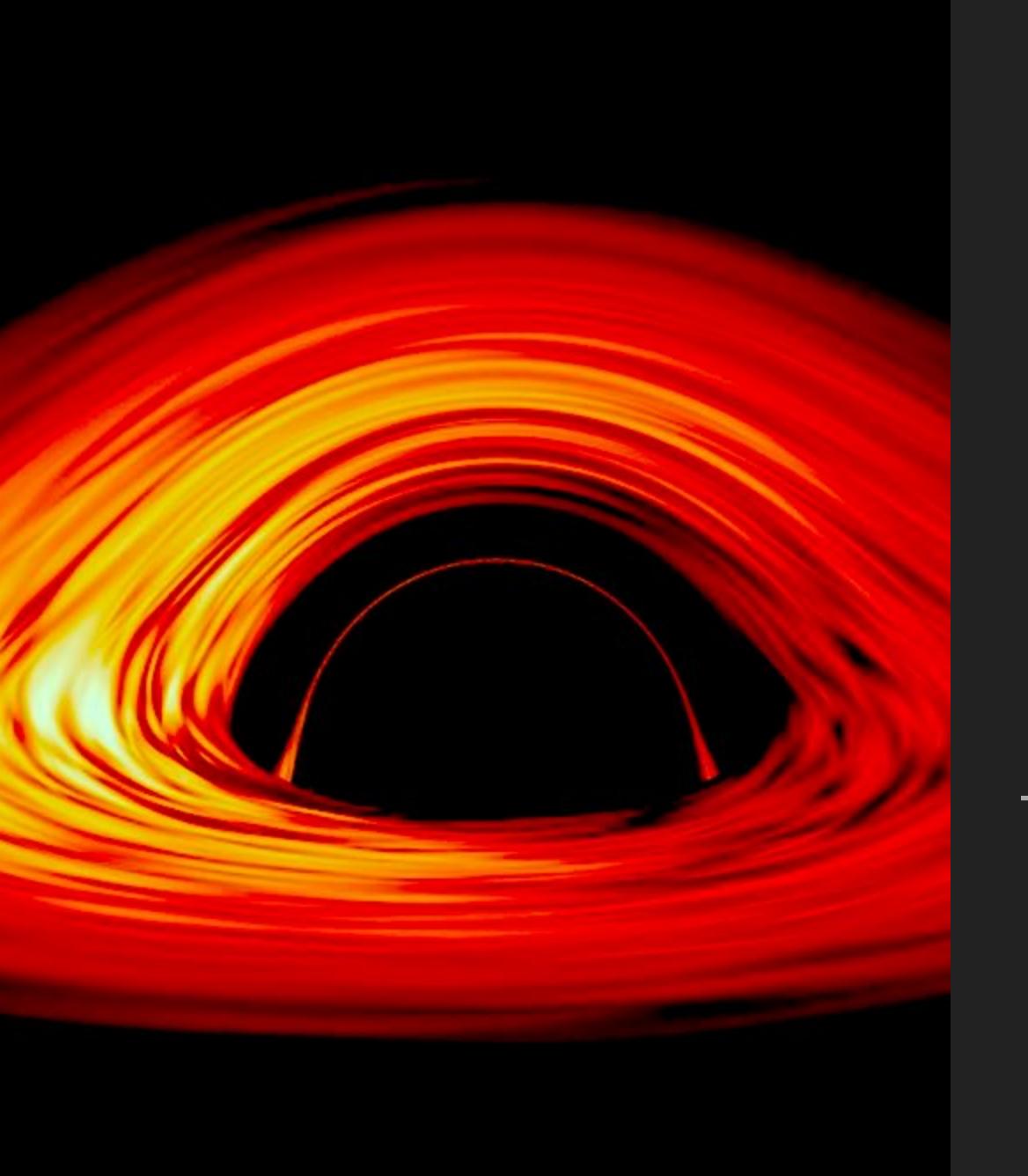
- Massive black holes (10⁵ 10⁷ M_o)
 - Their origin and growth
 - Merger physics, Electromagnetic counterparts
- Binary evolution & Milky Way structure
 - White dwarfs, neutron stars and black holes
 - Statistics and 3D positions
- Fundamental physics and testing GR
 - Black Hole physics
 - Gravitational wave properties
- Cosmology & Early Universe
 - Measure expansion
 - Signals from inflation, phase transitions, primordial black holes







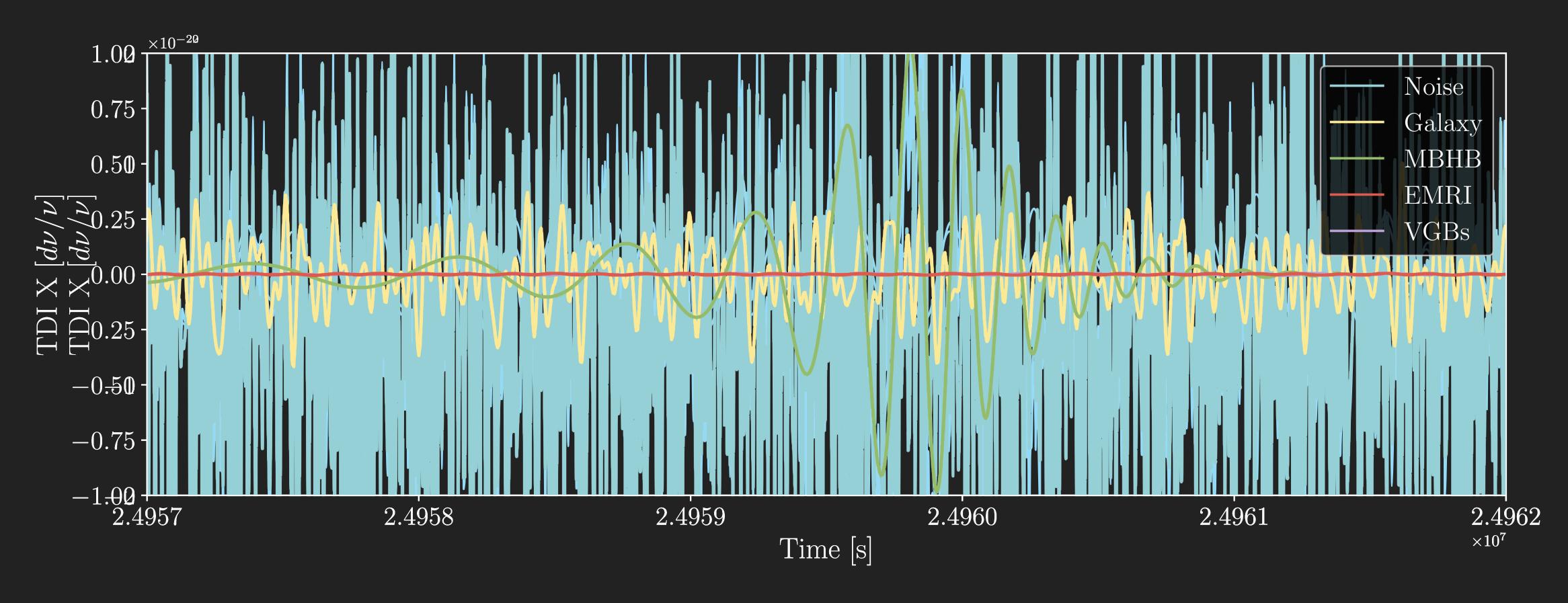




LISA

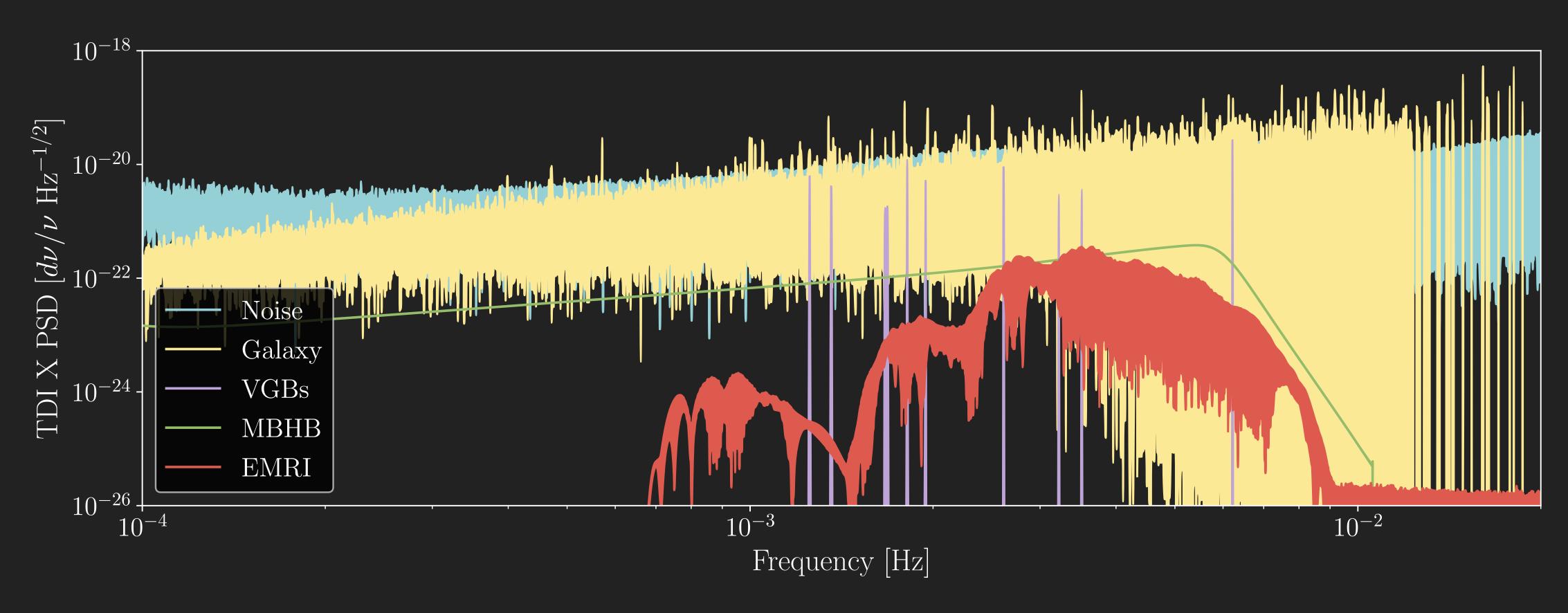
CHALLENGES

LISA DATA: MANY SIGNALS



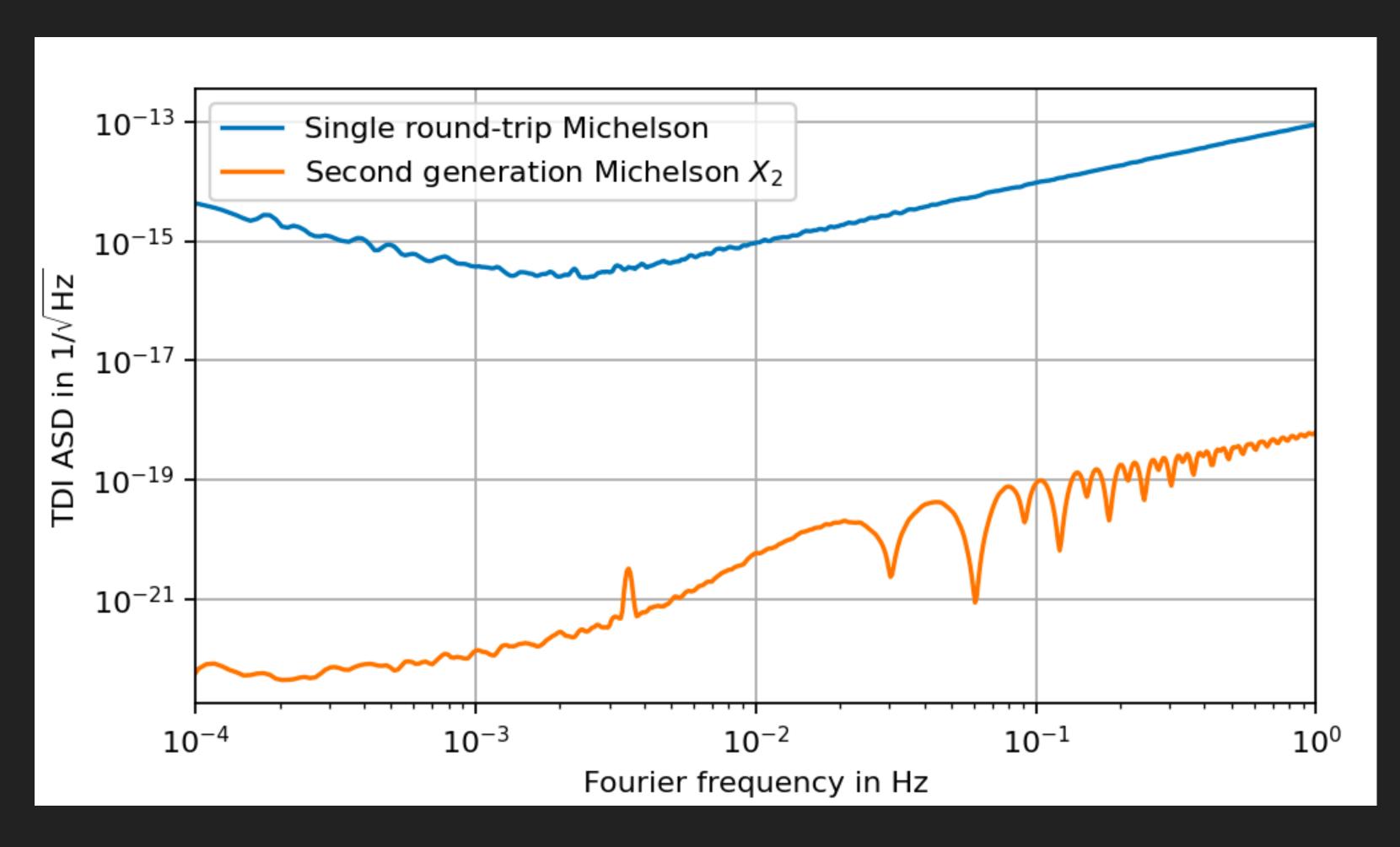
Verification GBs + EMRI + MBHB + Galaxy + Noise

LISA DATA: MANY SIGNALS



Verification GBs + EMRI + MBHB + Galaxy + Noise

SIGNALS BURIED IN LASER NOISE



TIME DELAY INTERFEROMETRY

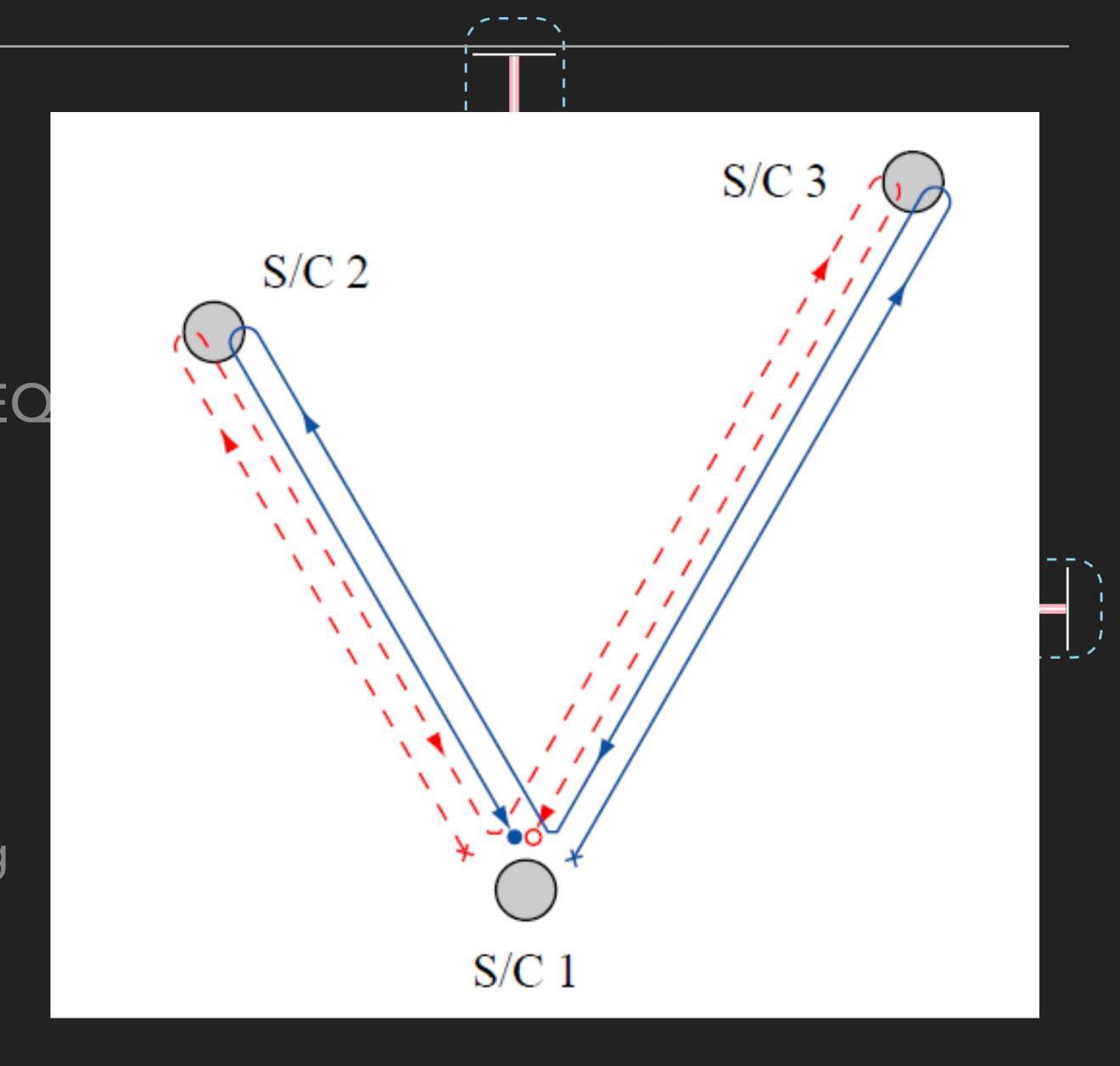
Why?

Laser Noise: 10^{-13}

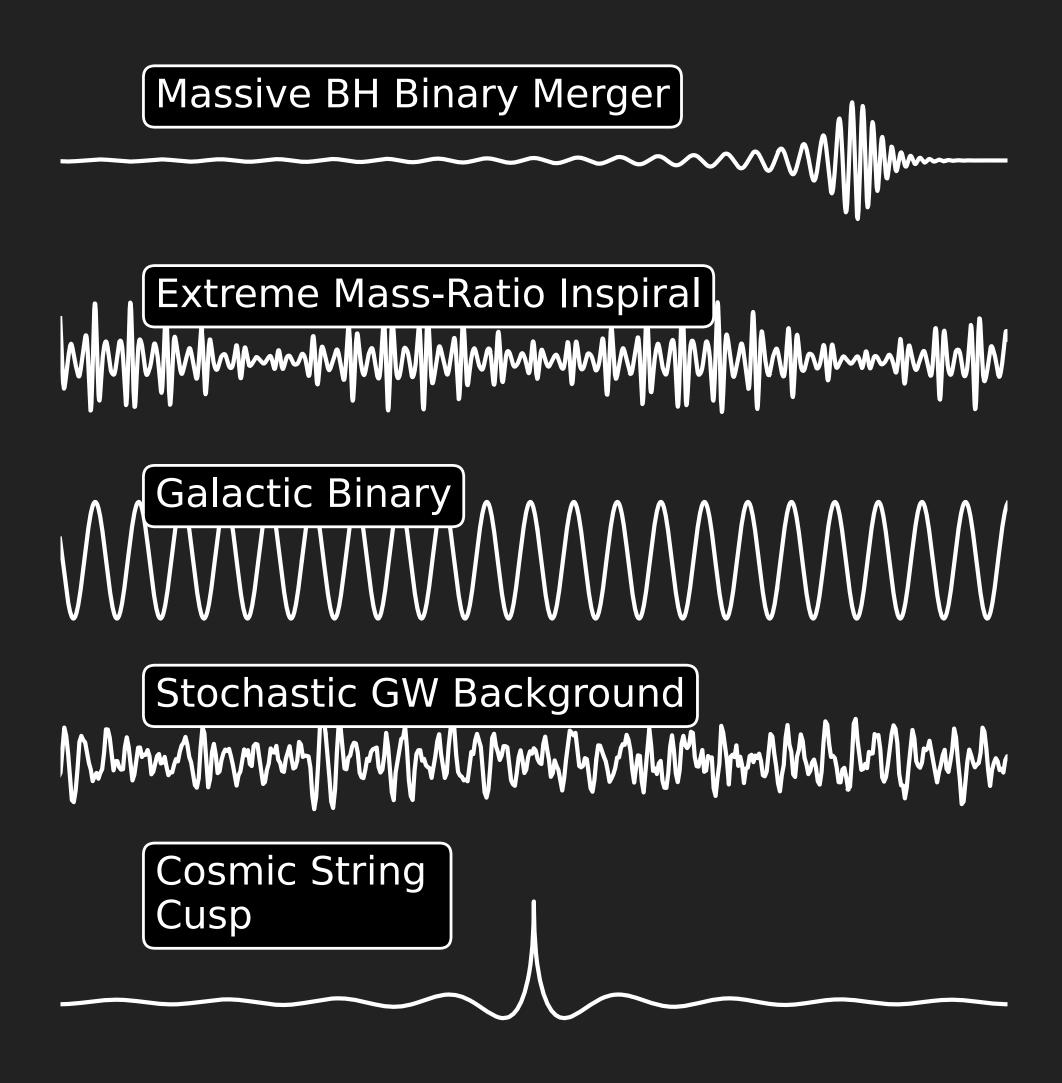
• Gravitational Waves: 10^{-20}

Solution:

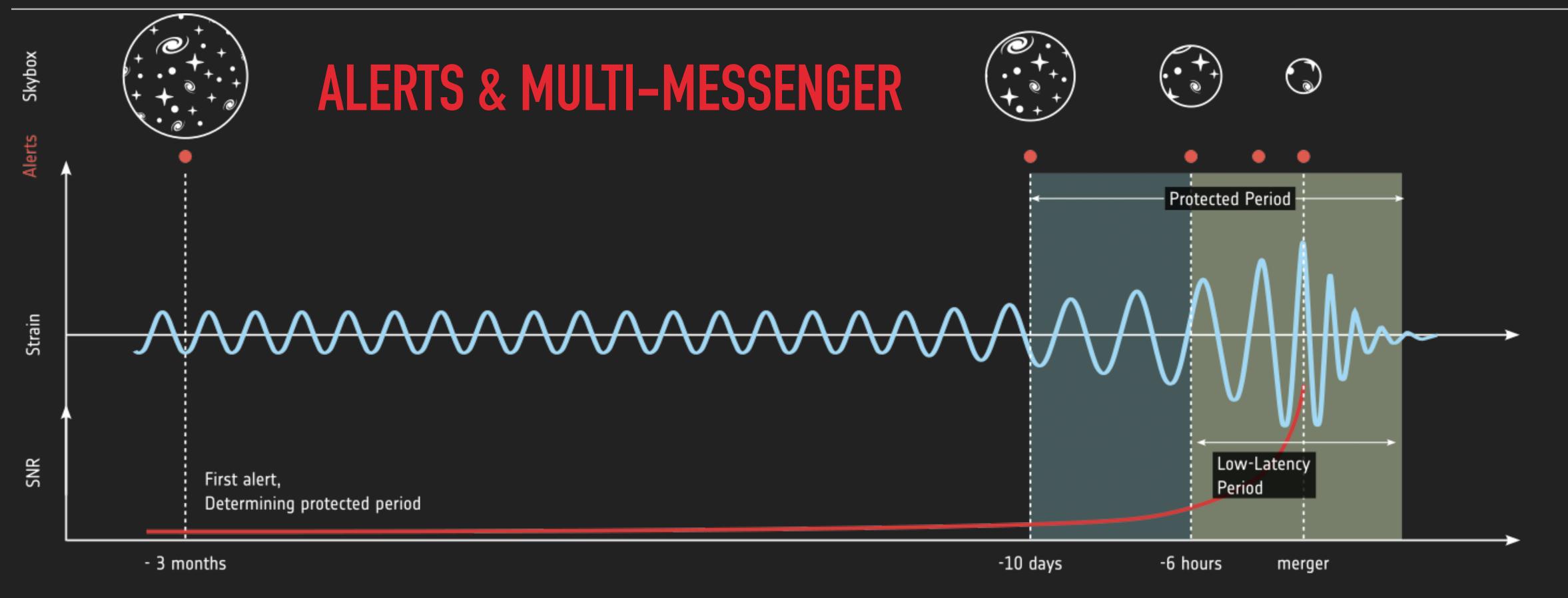
- Record data for each arm individually
- Delay and combine in post-processing
 - → Laser noise cancels out



WAVEFORMS CRUCIAL

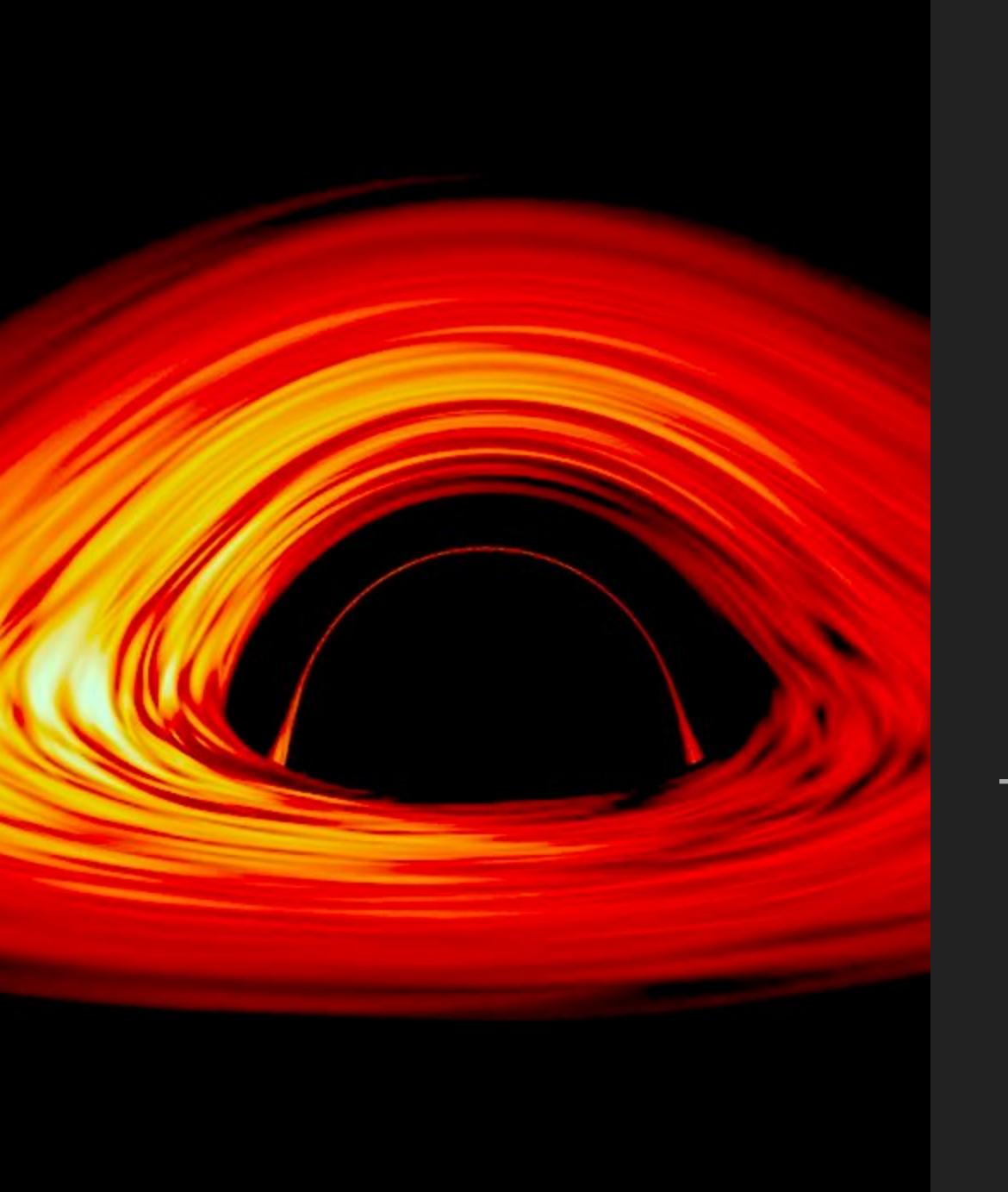


- Needed for Global Fit
- Status: good understanding, some serious development needed (EMRIs, IMRIs)
- Coordinated effort (Distributed Data Processing Center - DDPC, Consortium) to complete waveform models
- Different aspects
 - Accuracy for strong sources
 - Speed for discovery and low-latency searches



- Several low-latency alert pipelines
 - Alerts for new sources
 - Updates on known transients
- Realtime in 8/24 h of data exchange

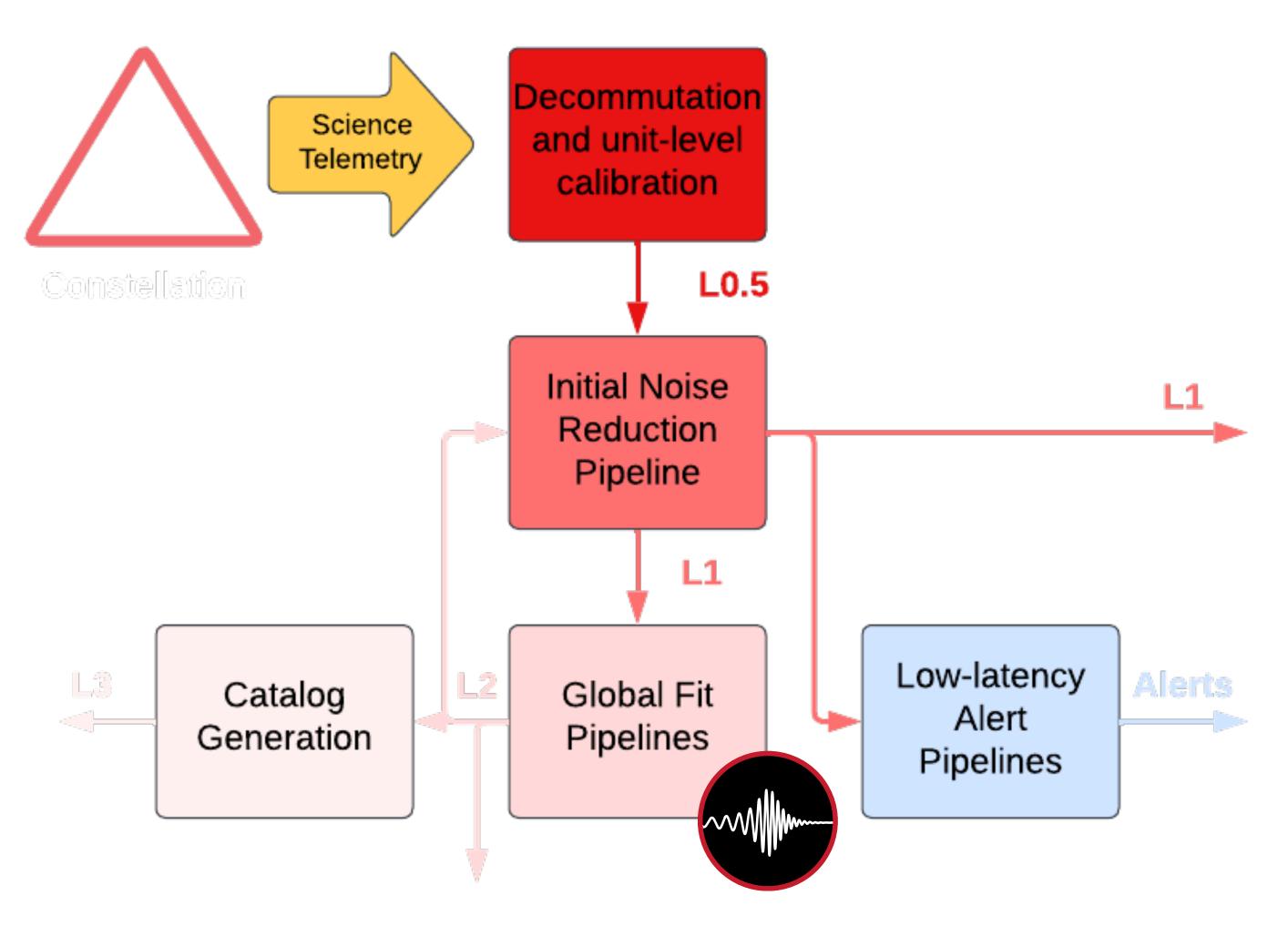
- Protected periods
- Multi-messenger will be done "online" with alerts but also "offline" for long lived sources (Galactic binaries, Stellar BH)



LISA

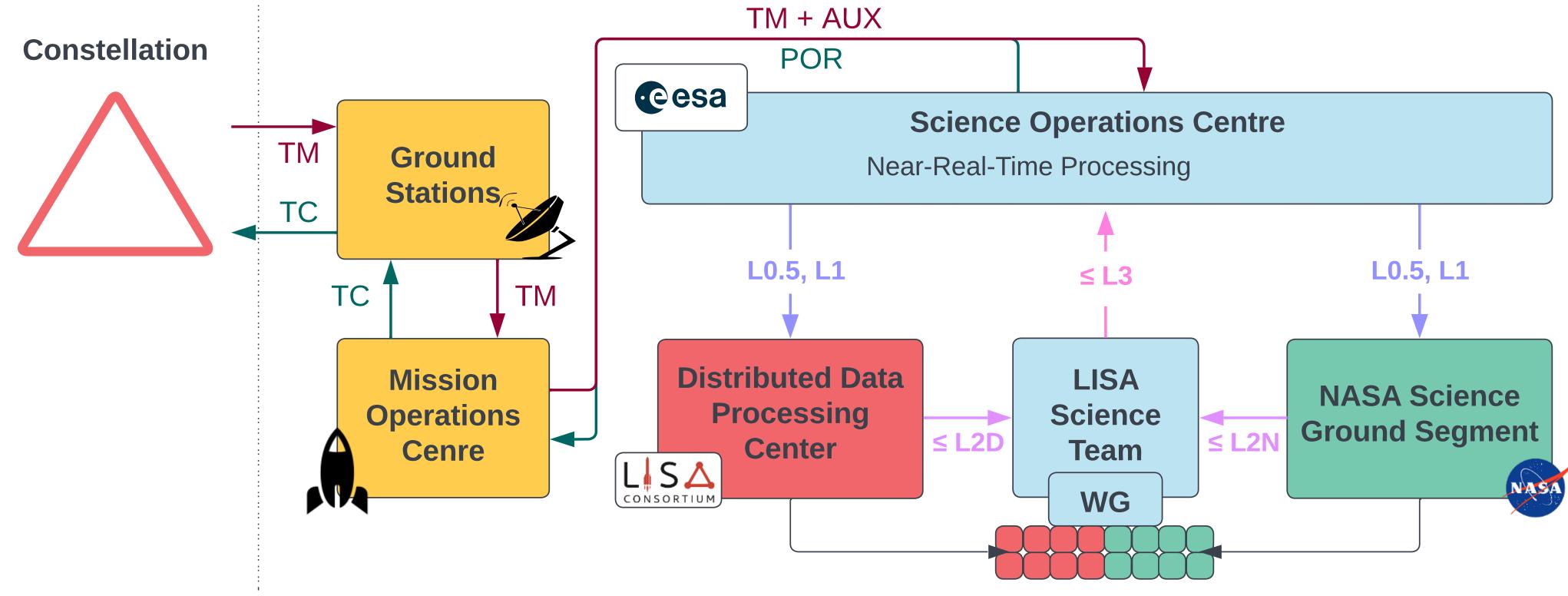
DATA PROCESSING

DISENTANGLING ALL SOURCES



- Main inputs: interferometric
 measurements + 3 armlengths
- Noise reduction: suppress laser noises and clock noises (Time Delay Interferometry)
- Global fits: iterative/simultaneous fitting for extracting a large number of overlapping sources
- Stochastic signals and residuals
- Detected sources combined to form catalogs

GROUND SEGMENT

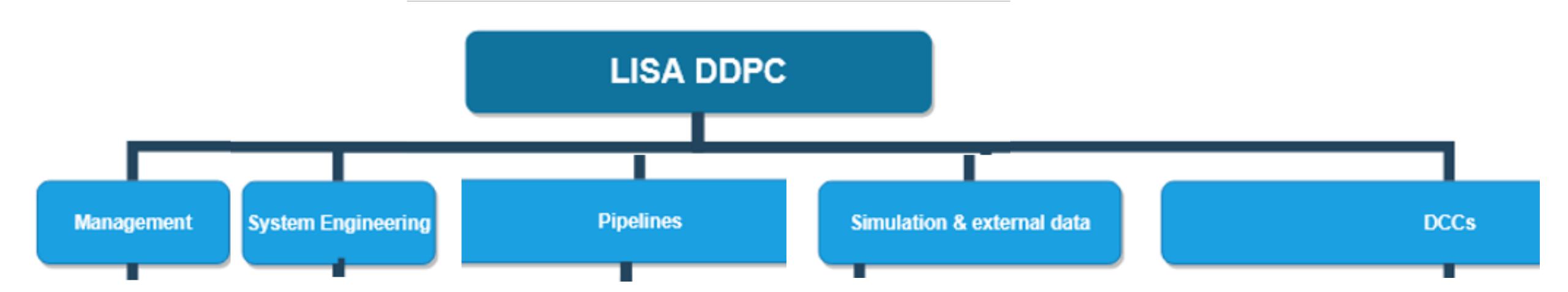


- Distributed Data Processing Center (DDPC), Europe
- NASA data processing

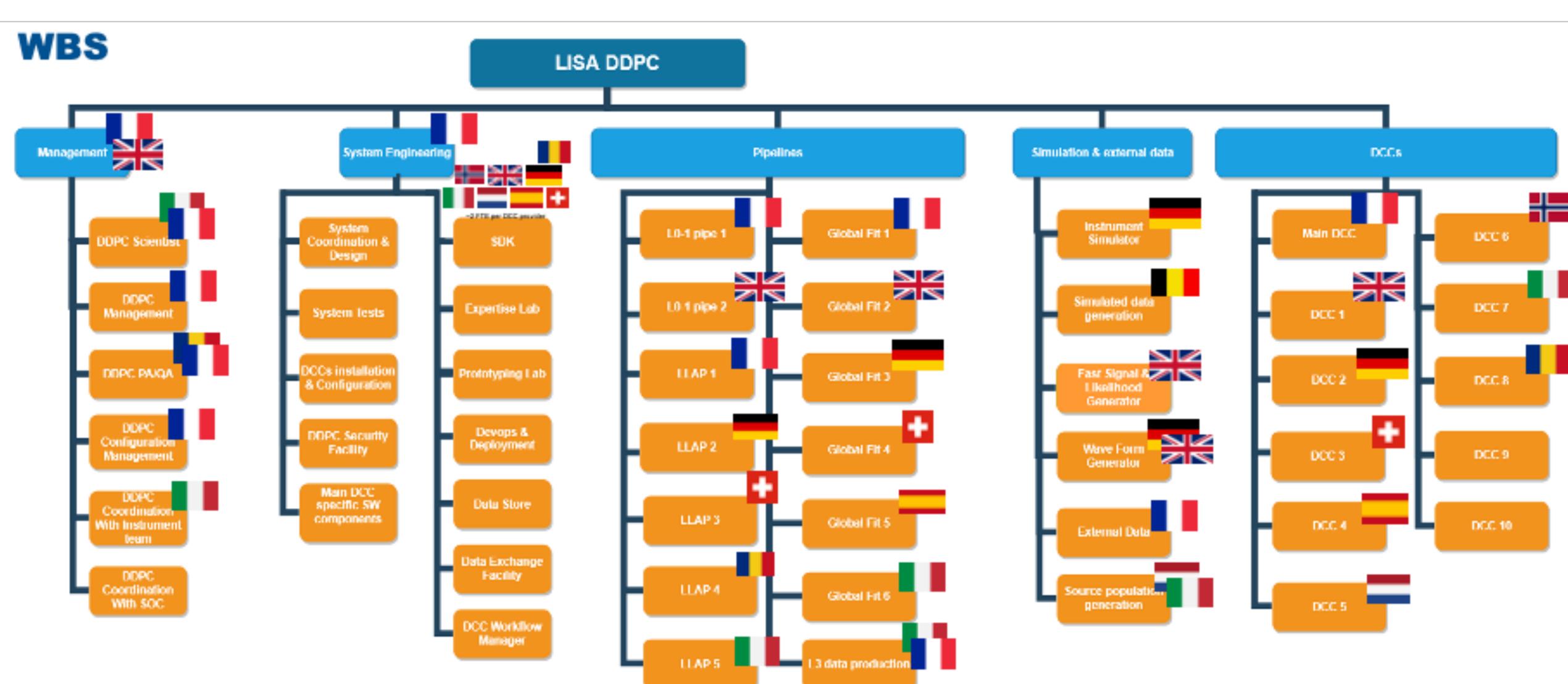
Several versions of L1, L2 data on different time scale

DDPC

- Responsibilities DDPC shared between ESA member states
- France in the lead
- NL will contribute
- Global structure



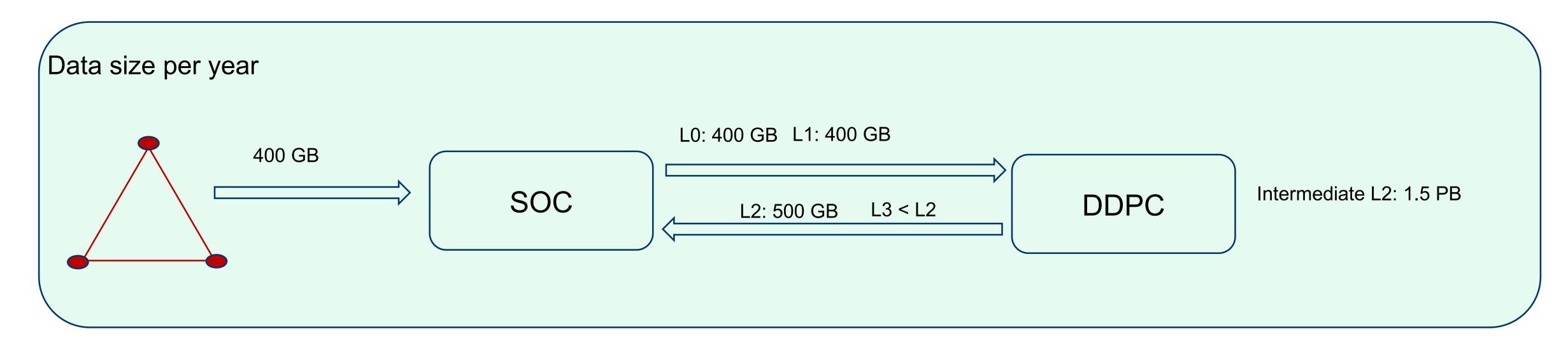
DDPC

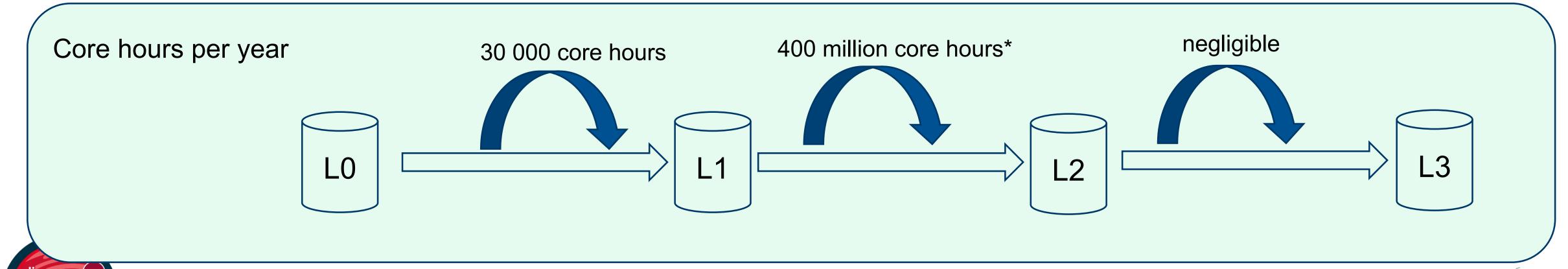


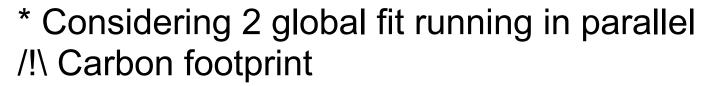
SCIENCE OBSERVATIONS: DATA PROCESSING, SCIENCE OPERATIONS AND DATA RELEASE



Some numbers



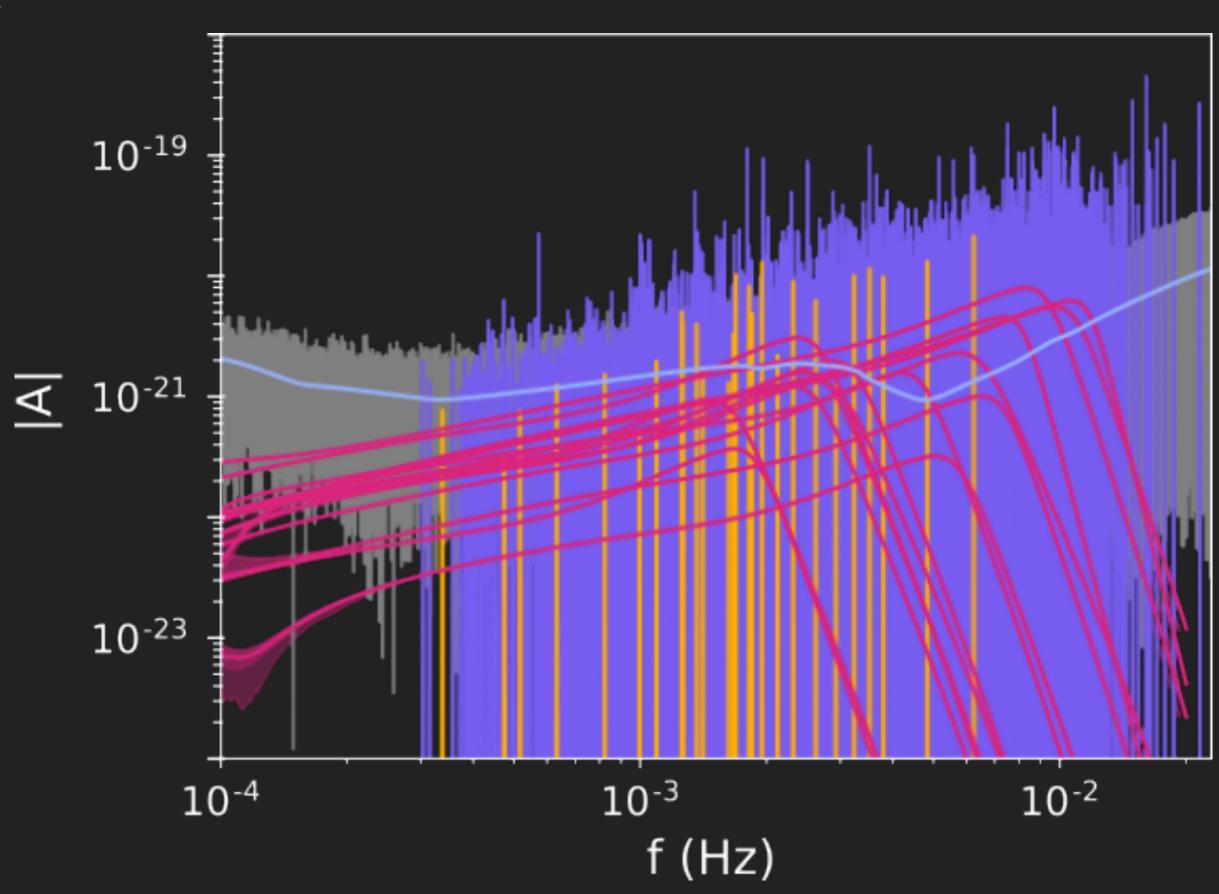






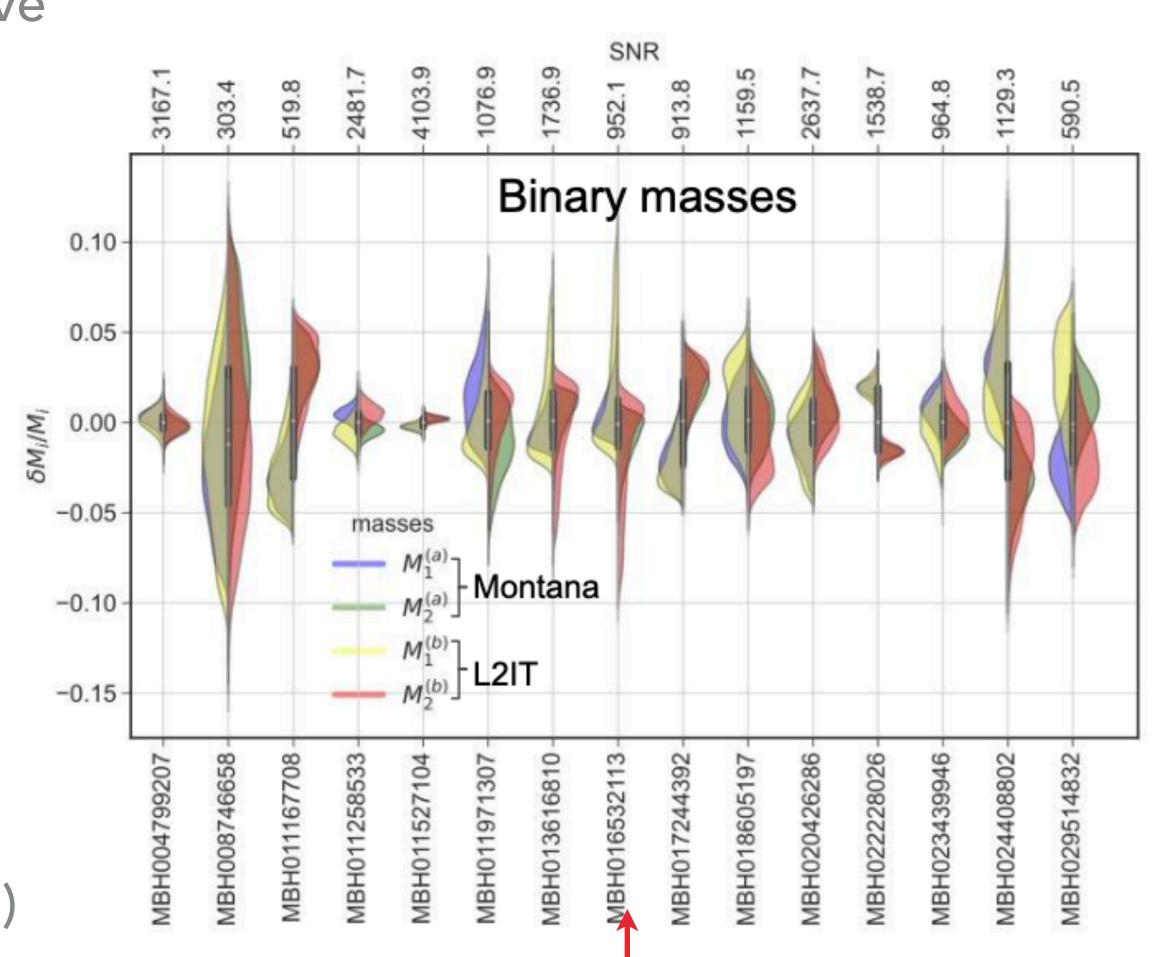
MAKING IT HAPPEN

- Distributed Data Processing Center (DDPC, Europe) & NASA Science Ground Segment
- Consolidated catalogs (L3) merged under ESA (PS/LST)
- Feasibility
 - LISA Data Challenges ongoing
 - Clear plan to final system
- Data will be made public, including tools and documentation



LISA DATA CHALLENGE

- Simulated datasets of increasing complexity to drive the development of LISA data analysis pipeline
- History:
 - 2005-2012: Mock LDC
 - (Re-)started in 2017
 - Future: via DDPC
- Example: Sangria, the last finalised challenge:
 1 year, all GBs + MBHBs + noises
 - Results for Galactic Binaries:
 - Injections with SNR > 8: 7800
 - Well recovered: 5000 to 6000
 - Partially recovered: 1000 to 2000 (up to 25%)
 - MBHBs



BIG DATA CHALLENGES

- Data processing
 - Synergy with other fields?
 - Role of community
- Data storage and management
 - ESA experience
 - Good plan, synergy/example possible
- Emerging technologies
 - ?

